

In Search of New Physics  
or  
Why We Need to Talk More!

Harrison B. Prosper

Florida State University

**WHEPP XIV, IIT Kanpur, India**

10 December, 2015

---

# Outline

- \* Introduction
- \* LHC Run I – A Random Stroll
- \* LHC Run II – Early Results
- \* Opportunities and Challenges
- \* Final Remarks

# Outline

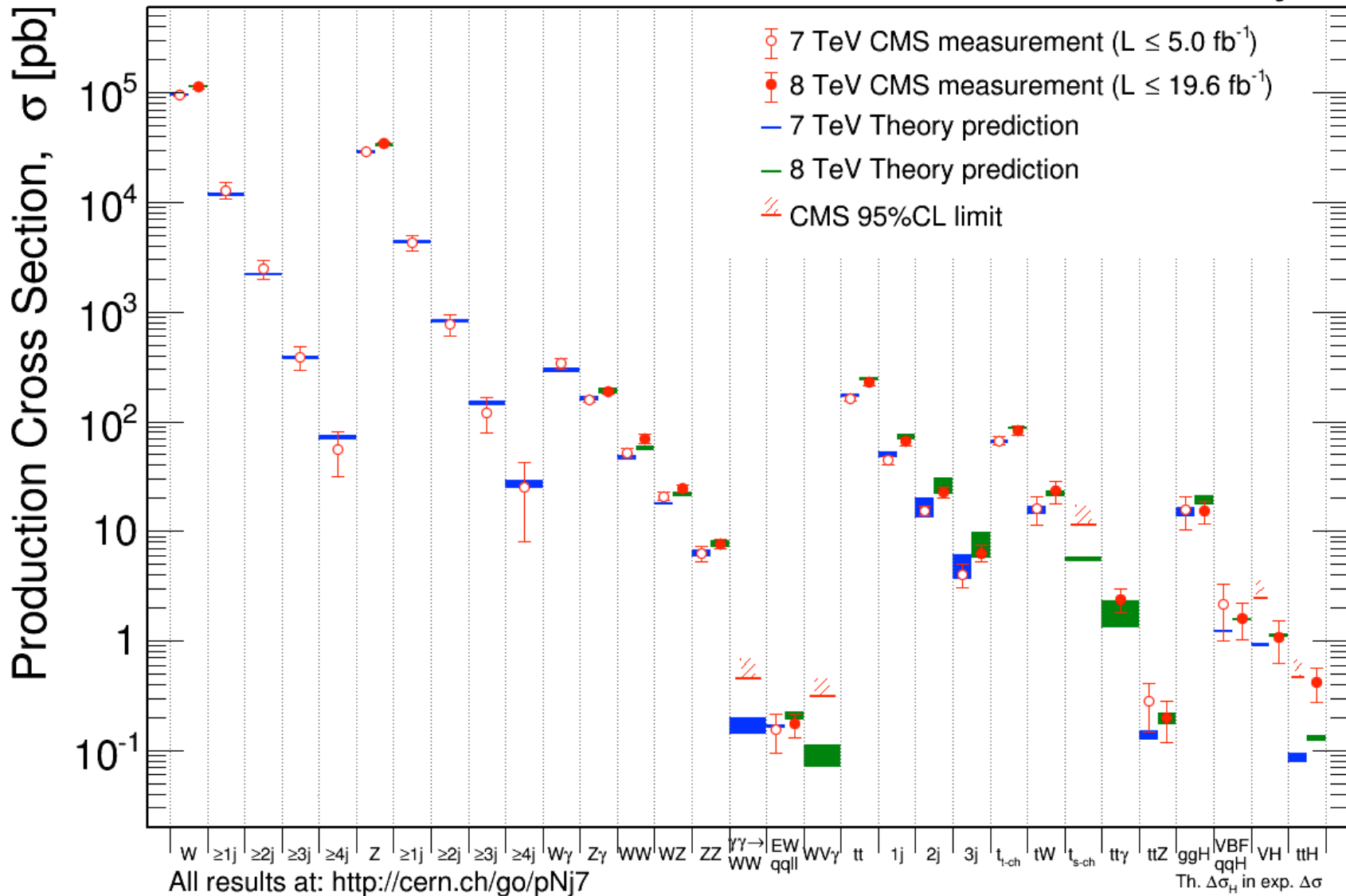
- \* **Introduction**
- \* LHC Run I – A Random Stroll
- \* LHC Run II – Early Results
- \* Opportunities and Challenges

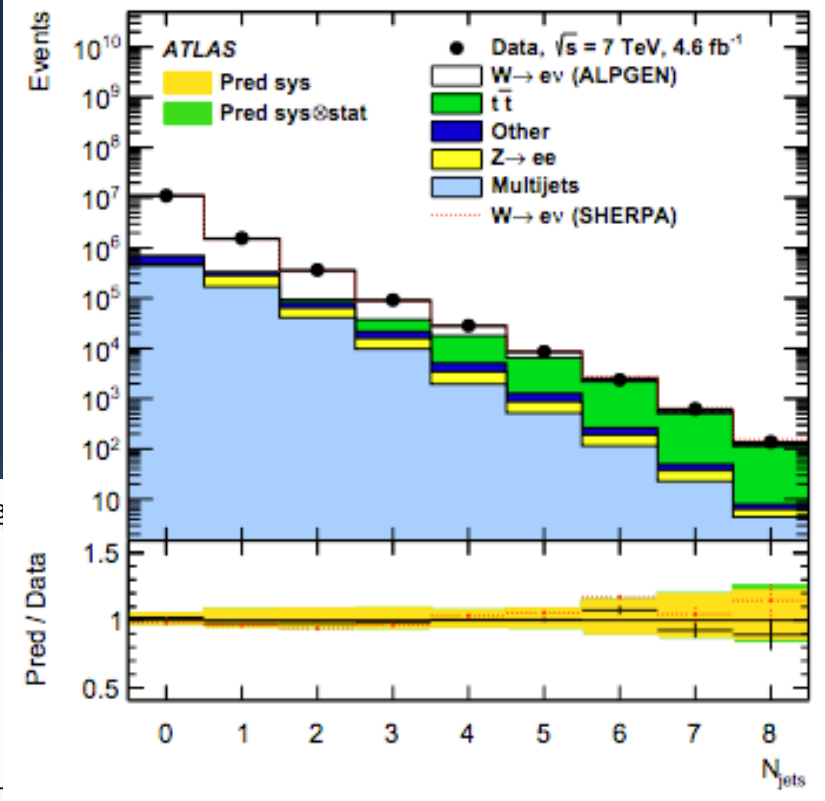
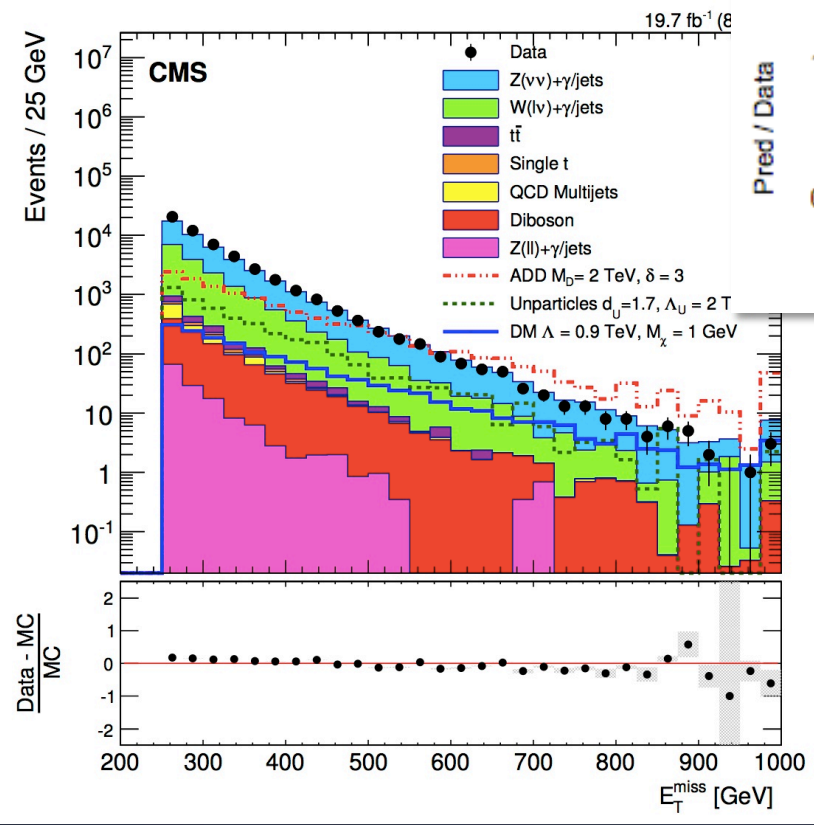
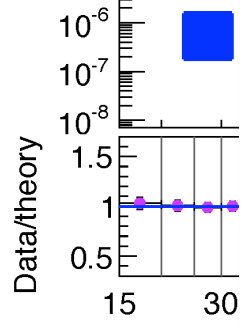
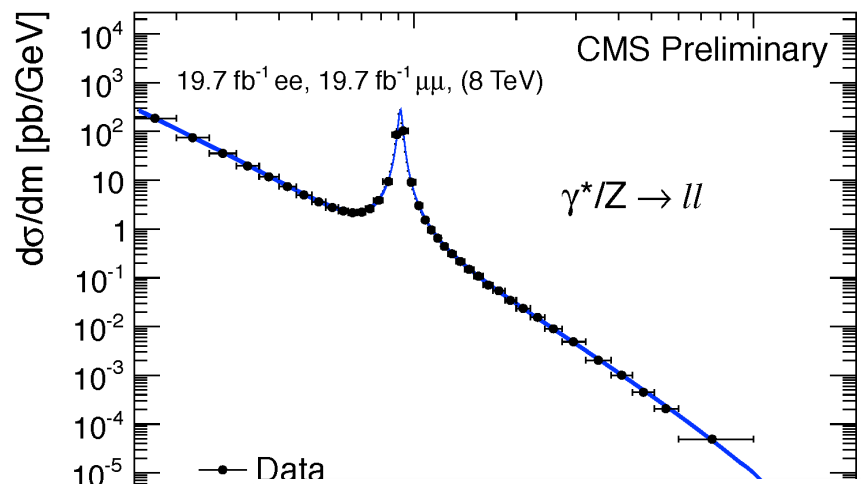
# Introduction

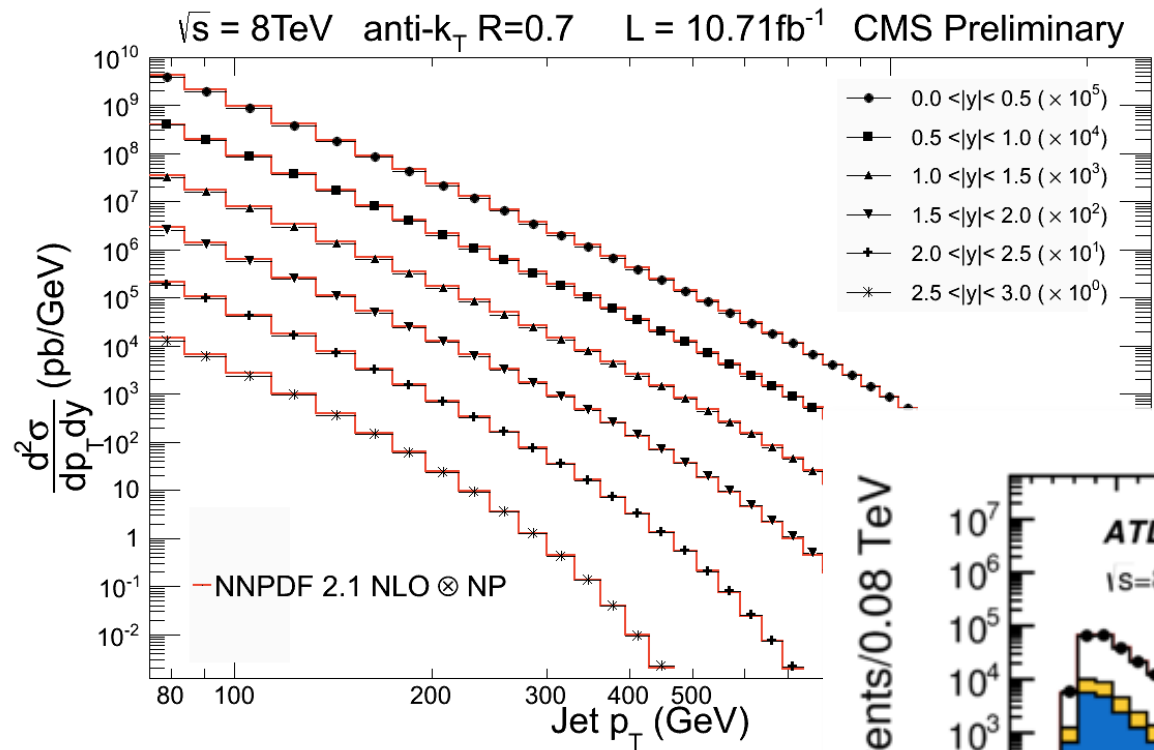
- \* The discovery of the Higgs boson at the LHC was a remarkable achievement.
- \* But, equally remarkable is the astonishing reaffirmation of the predictive power of the Standard Model.
- \* The disappointing corollary is the complete absence of compelling evidence of new physics at the LHC, so far.
- \* Nevertheless, there is still hope that we shall not be disappointed by the end of Run II.
- \* But, we could still be unlucky: finding no obvious resonances, but just a set of subtle spectral deviations.

Feb 2015

# CMS Preliminary

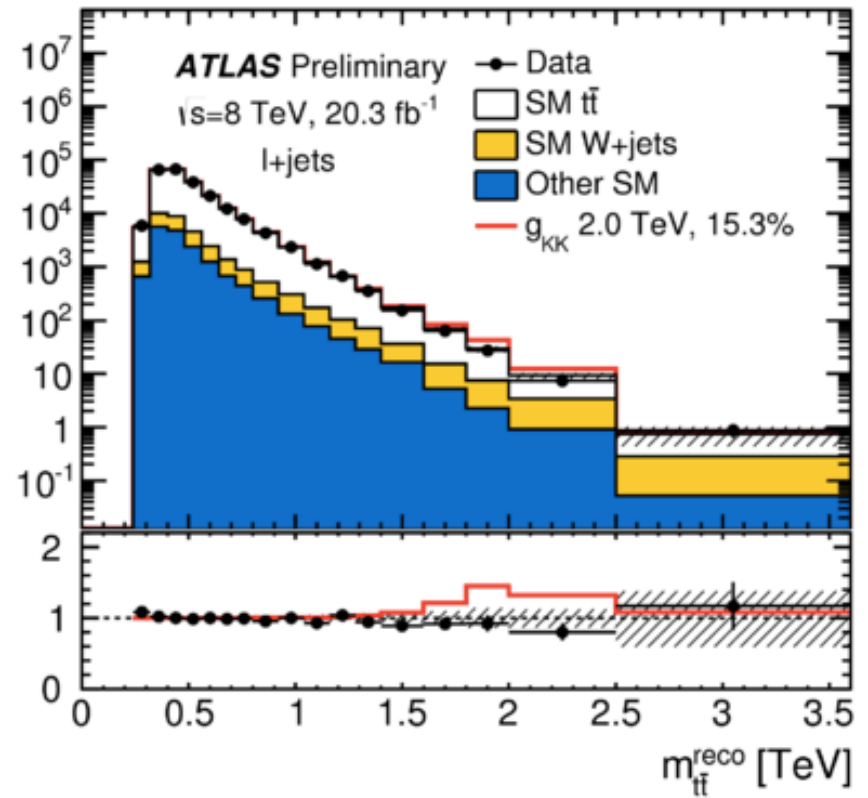






Events/0.08 TeV

Data/BG



# Introduction

- \* Ordinarily, the remarkable level of agreement between observations and predictions should be cause for celebration.
- \* But, that is not what we were hoping for in 2008!
- \* Run II, however, has only just started, and hope remains a rational stance.
- \* The fact that WHEPP is still going strong (XIV) suggests that many of you agree.



# Outline

- \* Introduction
- \* **LHC Run I – A Random Stroll**
- \* LHC Run II – Early Results
- \* The LHC and Beyond – Opportunities and Challenges



# LHC Run I – A Random Stroll

---

# Let “N” Thousand Flowers Bloom!

$$N \rightarrow \chi_0$$

Compositeness

Large extra dimensions

Heavy neutrinos

Di-boson resonances

Di-jet resonances

Braneworlds

Microscopic black holes

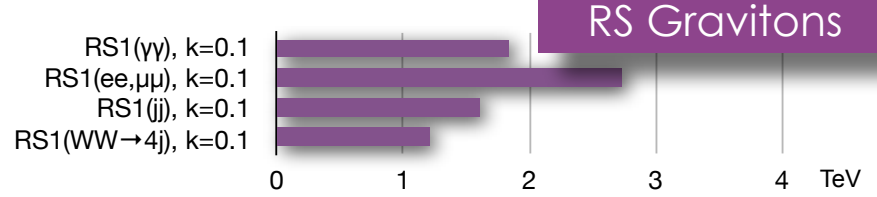
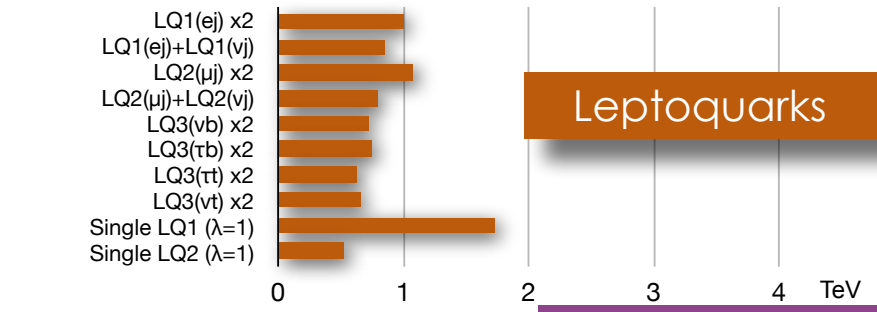
Supersymmetry

Leptoquarks

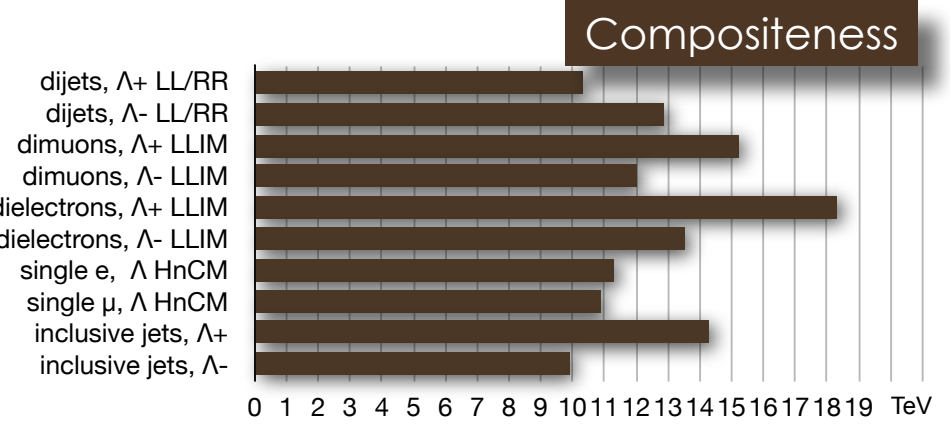
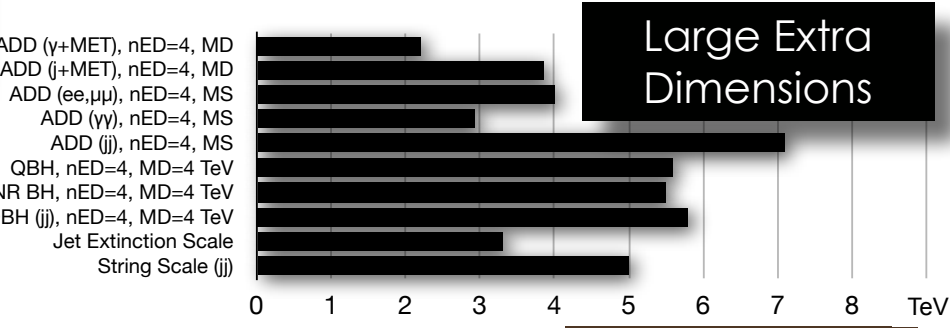
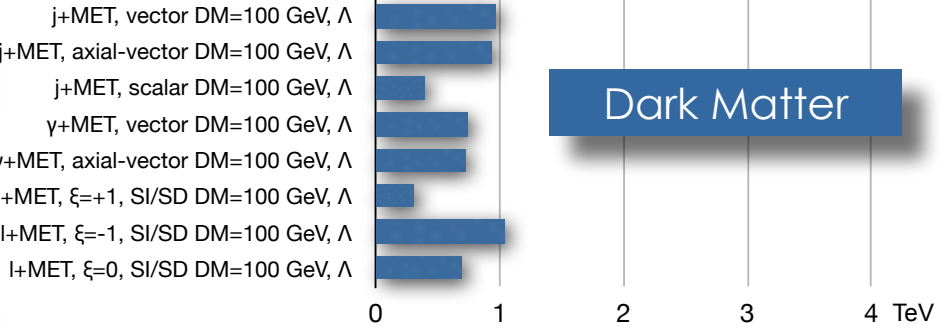
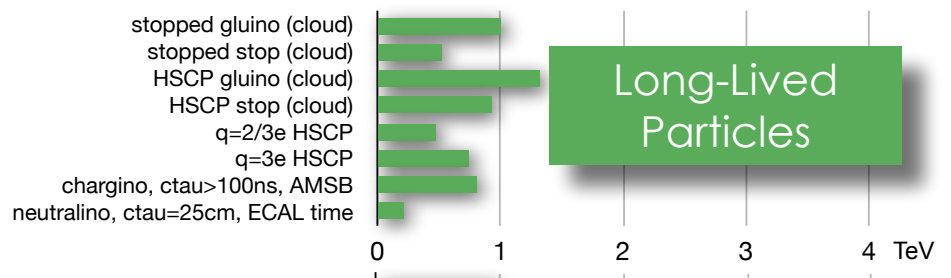
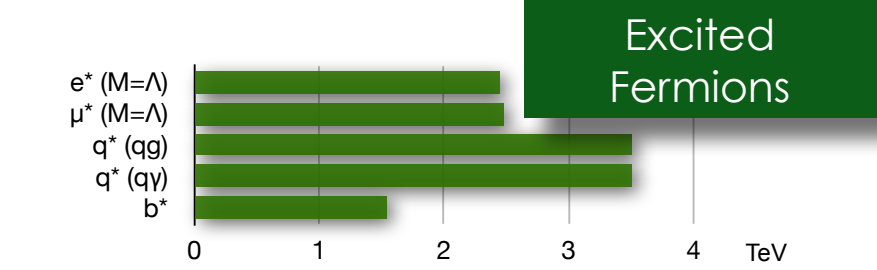
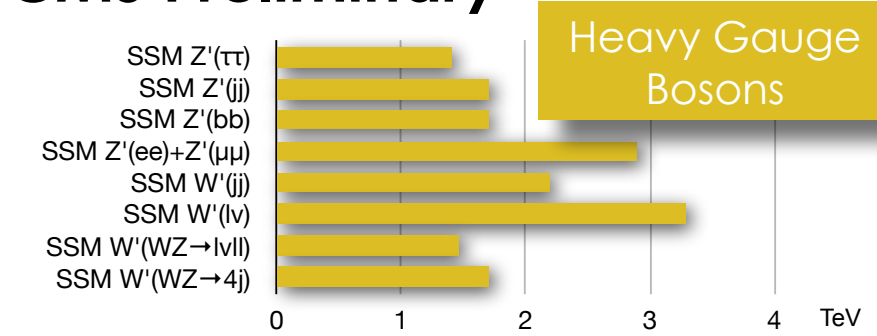
Walking, crawling, semi-comatose Technicolor

Strings

Excited fermions



**CMS Preliminary**



# ATLAS SUSY Searches\* - 95% CL Lower Limits

Status: July 2015

ATLAS Preliminary

$\sqrt{s} = 7, 8 \text{ TeV}$

Model	$e, \mu, \tau, \gamma$	Jets	$E_T^{\text{miss}}$	$\int \mathcal{L} d\mathcal{L} [\text{fb}^{-1}]$	Mass limit	$\sqrt{s} = 7 \text{ TeV}$	$\sqrt{s} = 8 \text{ TeV}$	Reference	
Inclusive Searches	MSUGRA/CMSSM	0-3 $e, \mu/1-2 \tau$	2-10 jets/3 $b$	Yes	20.3	$\tilde{q}, \tilde{g}$	1.8 TeV	$m(\tilde{q})=m(\tilde{g})$	1507.05525
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	$\tilde{q}$	850 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(1^{\text{st}} \text{ gen. } \tilde{q})=m(2^{\text{nd}} \text{ gen. } \tilde{q})$	1405.7875
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ (compressed)	mono-jet	1-3 jets	Yes	20.3	$\tilde{q}$	100-440 GeV	$m(\tilde{q})-m(\tilde{\chi}_1^0)<10 \text{ GeV}$	1507.05525
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q(\ell\ell/\ell\nu/\nu\nu)\tilde{\chi}_1^0$	2 $e, \mu$ (off-Z)	2 jets	Yes	20.3	$\tilde{q}$	780 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1503.03290
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	$\tilde{g}$	1.33 TeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1405.7875
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0 \rightarrow q\tilde{q}W^\pm\tilde{\chi}_1^0$	0-1 $e, \mu$	2-6 jets	Yes	20	$\tilde{g}$	1.26 TeV	$m(\tilde{\chi}_1^0)<300 \text{ GeV}, m(\tilde{\chi}^\pm)=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$	1507.05525
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell/\ell\nu/\nu\nu)\tilde{\chi}_1^0$	2 $e, \mu$	0-3 jets	-	20	$\tilde{g}$	1.32 TeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1501.03555
	GMSB ( $\tilde{\ell}$ NLSP)	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	20.3	$\tilde{g}$	1.6 TeV	$\tan\beta > 20$	1407.0603
	GGM (bino NLSP)	2 $\gamma$	-	Yes	20.3	$\tilde{g}$	1.29 TeV	$c\tau(\text{NLSP})<0.1 \text{ mm}$	1507.05493
	GGM (higgsino-bino NLSP)	$\gamma$	1 $b$	Yes	20.3	$\tilde{g}$	1.3 TeV	$m(\tilde{\chi}_1^0)<900 \text{ GeV}, c\tau(\text{NLSP})<0.1 \text{ mm}, \mu<0$	1507.05493
GGM (higgsino-bino NLSP)	$\gamma$	2 jets	Yes	20.3	$\tilde{g}$	1.25 TeV	$m(\tilde{\chi}_1^0)<850 \text{ GeV}, c\tau(\text{NLSP})<0.1 \text{ mm}, \mu>0$	1507.05493	
GGM (higgsino NLSP)	2 $e, \mu$ (Z)	2 jets	Yes	20.3	$\tilde{g}$	850 GeV	$m(\text{NLSP})>430 \text{ GeV}$	1503.03290	
Gravitino LSP	0	mono-jet	Yes	20.3	$\mu^{1/2}$ scale	865 GeV	$m(\tilde{G})>1.8 \times 10^{-4} \text{ eV}, m(\tilde{g})=m(\tilde{q})=1.5 \text{ TeV}$	1502.01518	
3 <sup>rd</sup> gen. $\tilde{g}$ med.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 $b$	Yes	20.1	$\tilde{g}$	1.25 TeV	$m(\tilde{\chi}_1^0)<400 \text{ GeV}$	1407.0600
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0	7-10 jets	Yes	20.3	$\tilde{g}$	1.1 TeV	$m(\tilde{\chi}_1^0)<350 \text{ GeV}$	1308.1841
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \mu\tilde{\mu}\tilde{\chi}_1^0$	0-1 $e, \mu$	3 $b$	Yes	20.1	$\tilde{g}$	1.34 TeV	$m(\tilde{\chi}_1^0)<400 \text{ GeV}$	1407.0600
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^+$	0-1 $e, \mu$	3 $b$	Yes	20.1	$\tilde{g}$	1.3 TeV	$m(\tilde{\chi}_1^0)<300 \text{ GeV}$	1407.0600
3 <sup>rd</sup> gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 $b$	Yes	20.1	$\tilde{b}_1$	100-620 GeV	$m(\tilde{\chi}_1^0)<90 \text{ GeV}$	1308.2631
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^\pm$	2 $e, \mu$ (SS)	0-3 $b$	Yes	20.3	$\tilde{b}_1$	275-440 GeV	$m(\tilde{\chi}_1^\pm)=2 m(\tilde{\chi}_1^0)$	1404.2500
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$	1-2 $e, \mu$	1-2 $b$	Yes	4.7/20.3	$\tilde{t}_1$	110-167 GeV	$m(\tilde{\chi}_1^\pm)=2 m(\tilde{\chi}_1^0), m(\tilde{\chi}_1^0)=55 \text{ GeV}$	1209.2102, 1407.0583
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $t\tilde{\chi}_1^0$	0-2 $e, \mu$	0-2 jets/1-2 $b$	Yes	20.3	$\tilde{t}_1$	90-191 GeV	$m(\tilde{\chi}_1^0)=1 \text{ GeV}$	1506.08616
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet/ $c$ -tag	Yes	20.3	$\tilde{t}_1$	90-240 GeV	$m(\tilde{t}_1)-m(\tilde{\chi}_1^0)<85 \text{ GeV}$	1407.0608
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 $e, \mu$ (Z)	1 $b$	Yes	20.3	$\tilde{t}_1$	150-580 GeV	$m(\tilde{\chi}_1^0)>150 \text{ GeV}$	1403.5222
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 $e, \mu$ (Z)	1 $b$	Yes	20.3	$\tilde{t}_2$	290-600 GeV	$m(\tilde{\chi}_1^0)<200 \text{ GeV}$	1403.5222	
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 $e, \mu$	0	Yes	20.3	$\tilde{\ell}$	90-325 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1403.5294
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\ell}\nu(\tilde{\nu})$	2 $e, \mu$	0	Yes	20.3	$\tilde{\chi}_1^\pm$	140-465 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1403.5294
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\nu}\nu(\tilde{\nu}\tilde{\nu})$	2 $\tau$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	100-350 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(\tilde{\tau}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1407.0350
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_L\nu\tilde{\ell}_L(\tilde{\nu}\tilde{\nu}), \tilde{\ell}\tilde{\nu}\tilde{\ell}_L(\tilde{\nu}\tilde{\nu})$	3 $e, \mu$	0	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	700 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1402.7029
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 Z\tilde{\chi}_1^0$	2-3 $e, \mu$	0-2 jets	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	420 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1403.5294, 1402.7029
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 h\tilde{\chi}_1^0, h \rightarrow b\tilde{b}/WW/\tau\tau/\gamma\gamma$	$e, \mu, \gamma$	0-2 $b$	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	250 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, \text{ sleptons decoupled}$	1501.07110
	$\tilde{\chi}_2^0\tilde{\chi}_3^0, \tilde{\chi}_2^0 \rightarrow \tilde{\ell}_R\ell$	4 $e, \mu$	0	Yes	20.3	$\tilde{\chi}_2^0, \tilde{\chi}_3^0$	620 GeV	$m(\tilde{\chi}_2^0)=m(\tilde{\chi}_3^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_2^0)+m(\tilde{\chi}_1^0))$	1405.5086
	GGM (wino NLSP) weak prod.	1 $e, \mu + \gamma$	-	Yes	20.3	$\tilde{W}$	124-361 GeV	$c\tau<1 \text{ mm}$	1507.05493
Long-lived particles	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	20.3	$\tilde{\chi}_1^\pm$	270 GeV	$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)\sim 160 \text{ MeV}, \tau(\tilde{\chi}_1^\pm)=0.2 \text{ ns}$	1310.3675
	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$	dE/dx trk	-	Yes	18.4	$\tilde{\chi}_1^\pm$	482 GeV	$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)\sim 160 \text{ MeV}, \tau(\tilde{\chi}_1^\pm)<15 \text{ ns}$	1506.05332
	Stable, stopped $\tilde{g}$ R-hadron	0	1-5 jets	Yes	27.9	$\tilde{g}$	832 GeV	$m(\tilde{\chi}_1^0)=100 \text{ GeV}, 10 \mu\text{s}<\tau(\tilde{g})<1000 \text{ s}$	1310.6584
	Stable $\tilde{g}$ R-hadron	trk	-	-	19.1	$\tilde{g}$	1.27 TeV	-	1411.6795
	GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu})+\tau(e, \mu)$	1-2 $\mu$	-	-	19.1	$\tilde{\chi}_1^0$	537 GeV	$10<\tan\beta<50$	1411.6795
	GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$ , long-lived $\tilde{\chi}_1^0$	2 $\gamma$	-	Yes	20.3	$\tilde{\chi}_1^0$	435 GeV	$2<\tau(\tilde{\chi}_1^0)<3 \text{ ns}, \text{ SPS8 model}$	1409.5542
	$\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow e\tilde{e}\nu/\mu\tilde{\mu}\nu$	displ. $e\tilde{e}/\mu\tilde{\mu}\nu$	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$7<c\tau(\tilde{\chi}_1^0)<740 \text{ mm}, m(\tilde{g})=1.3 \text{ TeV}$	1504.05162
	GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$	displ. vtx + jets	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$6<c\tau(\tilde{\chi}_1^0)<480 \text{ mm}, m(\tilde{g})=1.1 \text{ TeV}$	1504.05162
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\tau\mu$	$e\mu, e\tau, \mu\tau$	-	-	20.3	$\tilde{\nu}_\tau$	1.7 TeV	$\lambda'_{311}=0.11, \lambda'_{132/133/233}=0.07$	1503.04430
	Bilinear RPV CMSSM	2 $e, \mu$ (SS)	0-3 $b$	Yes	20.3	$\tilde{q}, \tilde{g}$	1.35 TeV	$m(\tilde{q})=m(\tilde{g}), c\tau_{\text{LSP}}<1 \text{ mm}$	1404.2500
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow e\tilde{\nu}_\mu, e\mu\tilde{\nu}_e$	4 $e, \mu$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	750 GeV	$m(\tilde{\chi}_1^0)>0.2 \times m(\tilde{\chi}_1^\pm), \lambda_{121} \neq 0$	1405.5086
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tilde{\nu}_e, e\tau\tilde{\nu}_\tau$	3 $e, \mu + \tau$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	450 GeV	$m(\tilde{\chi}_1^0)>0.2 \times m(\tilde{\chi}_1^\pm), \lambda_{133} \neq 0$	1405.5086
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}q$	0	6-7 jets	-	20.3	$\tilde{g}$	917 GeV	$\text{BR}(t)=\text{BR}(b)=\text{BR}(c)=0\%$	1502.05686
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow q\tilde{q}q$	0	6-7 jets	-	20.3	$\tilde{g}$	870 GeV	$m(\tilde{\chi}_1^0)=600 \text{ GeV}$	1502.05686
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t_1 t_1, \tilde{t}_1 \rightarrow bs$	2 $e, \mu$ (SS)	0-3 $b$	Yes	20.3	$\tilde{g}$	850 GeV	-	1404.250
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 $b$	-	20.3	$\tilde{t}_1$	100-308 GeV	-	ATLAS-CONF-2015-026
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\ell}$	2 $e, \mu$	2 $b$	-	20.3	$\tilde{t}_1$	0.4-1.0 TeV	$\text{BR}(\tilde{t}_1 \rightarrow b\ell/\mu)>20\%$	ATLAS-CONF-2015-015	
Other	Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$	0	2 $c$	Yes	20.3	$\tilde{c}$	490 GeV	$m(\tilde{\chi}_1^0)<200 \text{ GeV}$	1501.01325

$10^{-1}$

1

Mass scale [TeV]

\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus  $1\sigma$  theoretical signal cross section uncertainty.

# Searching for Resonances

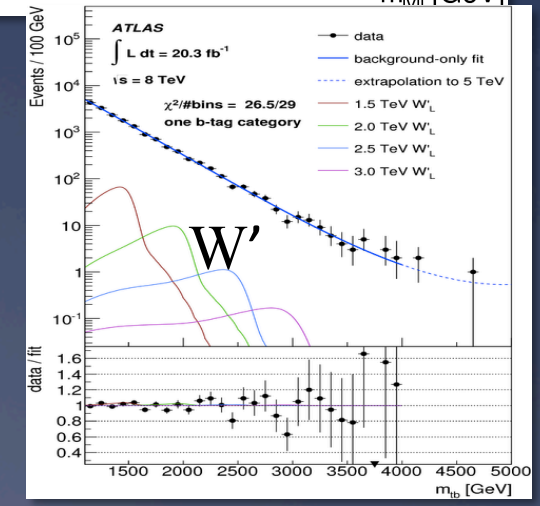
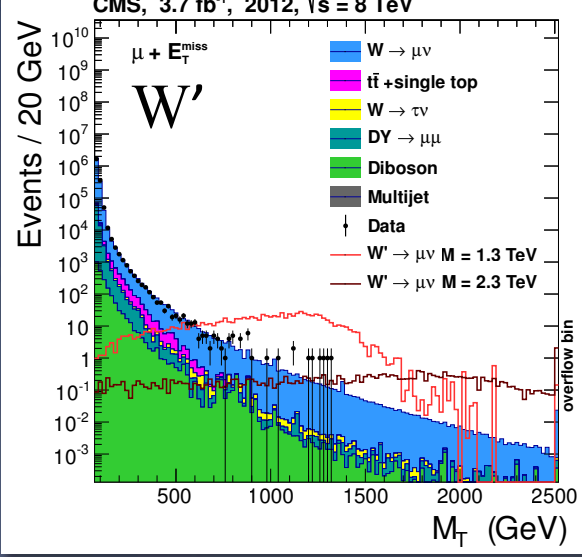
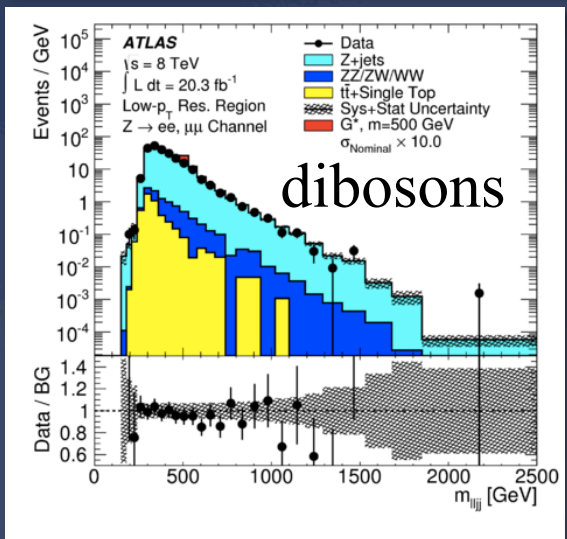
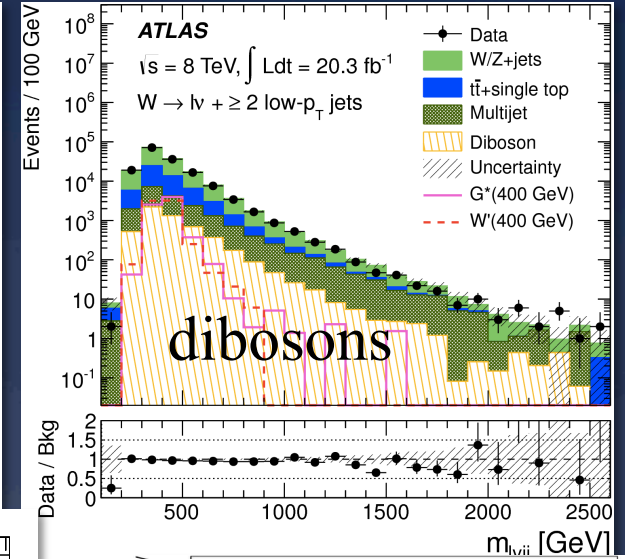
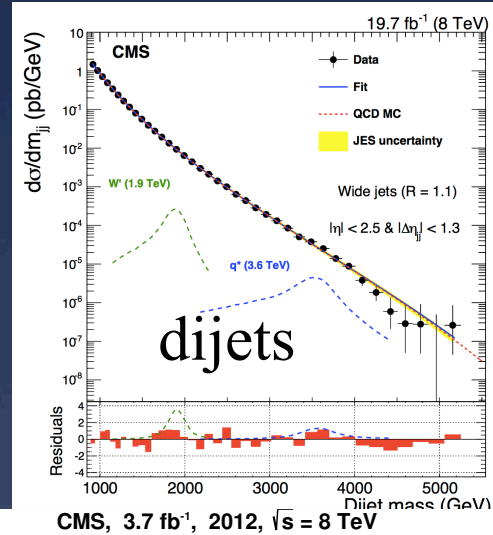
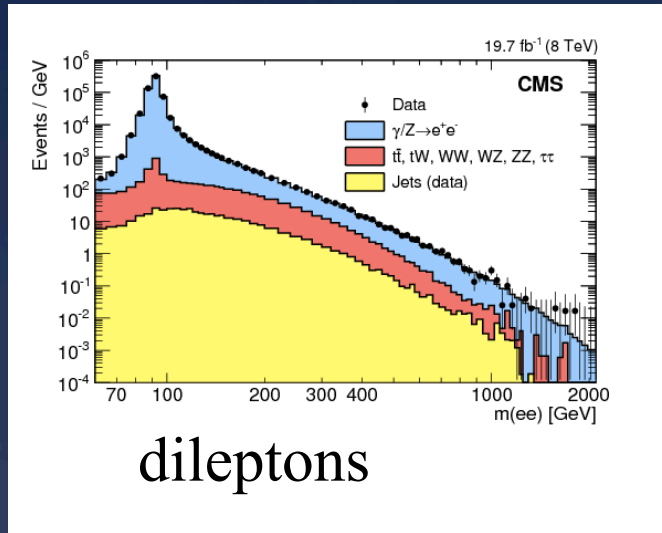
Resonances, such as

- \* excited quarks
- \*  $W$  and  $Z$  primes
- \* Randall-Sundrum gravitons
- \* Axiguons,

to name but a few, are a generic prediction of many models of beyond the SM physics.

A systematic search for resonances should be a priority.

# Searching for Resonances



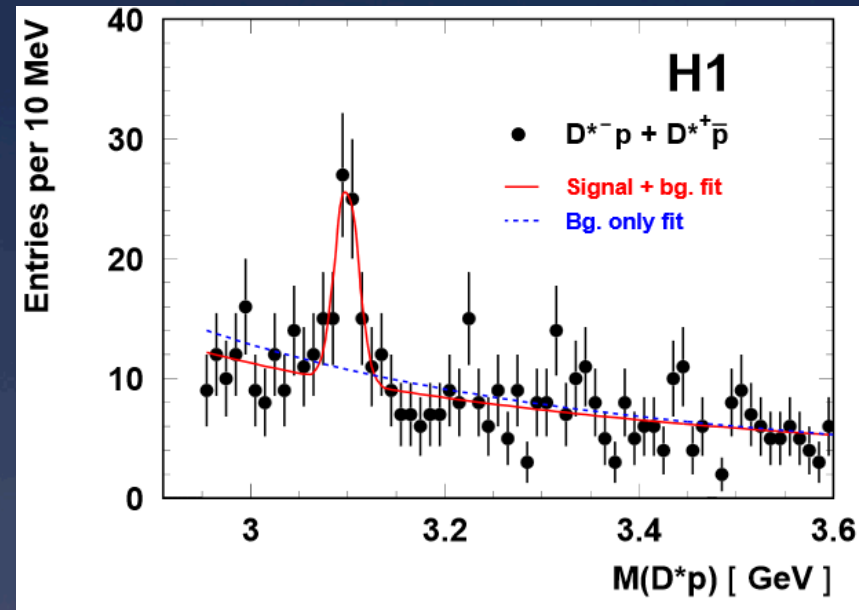
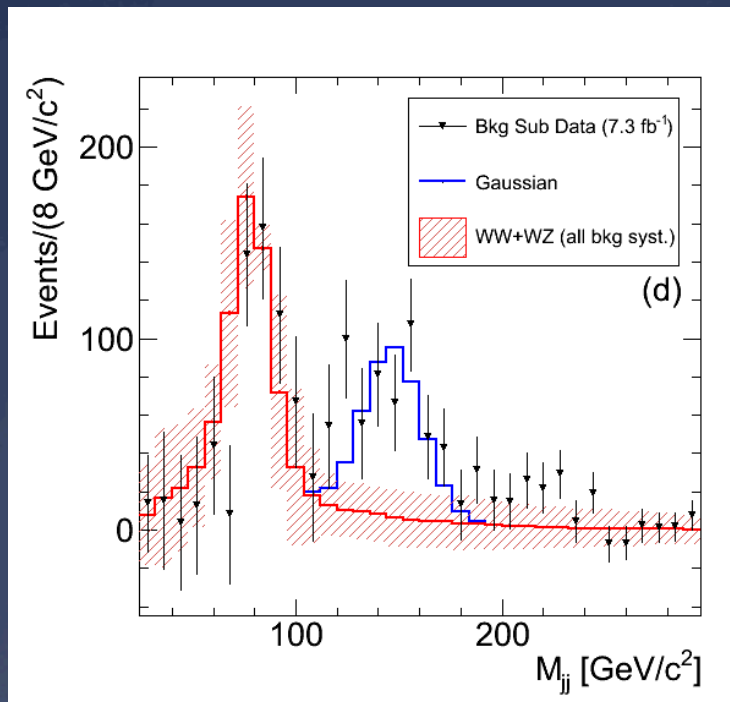
# A (Cautionary) Tale of Resonances

- \* Today, no discovery in our field is more convincing than observing a peak, especially if narrow, on top of a smooth monotonic background.
  
- \* But, we should tread carefully...



# Resonances That Aren't!

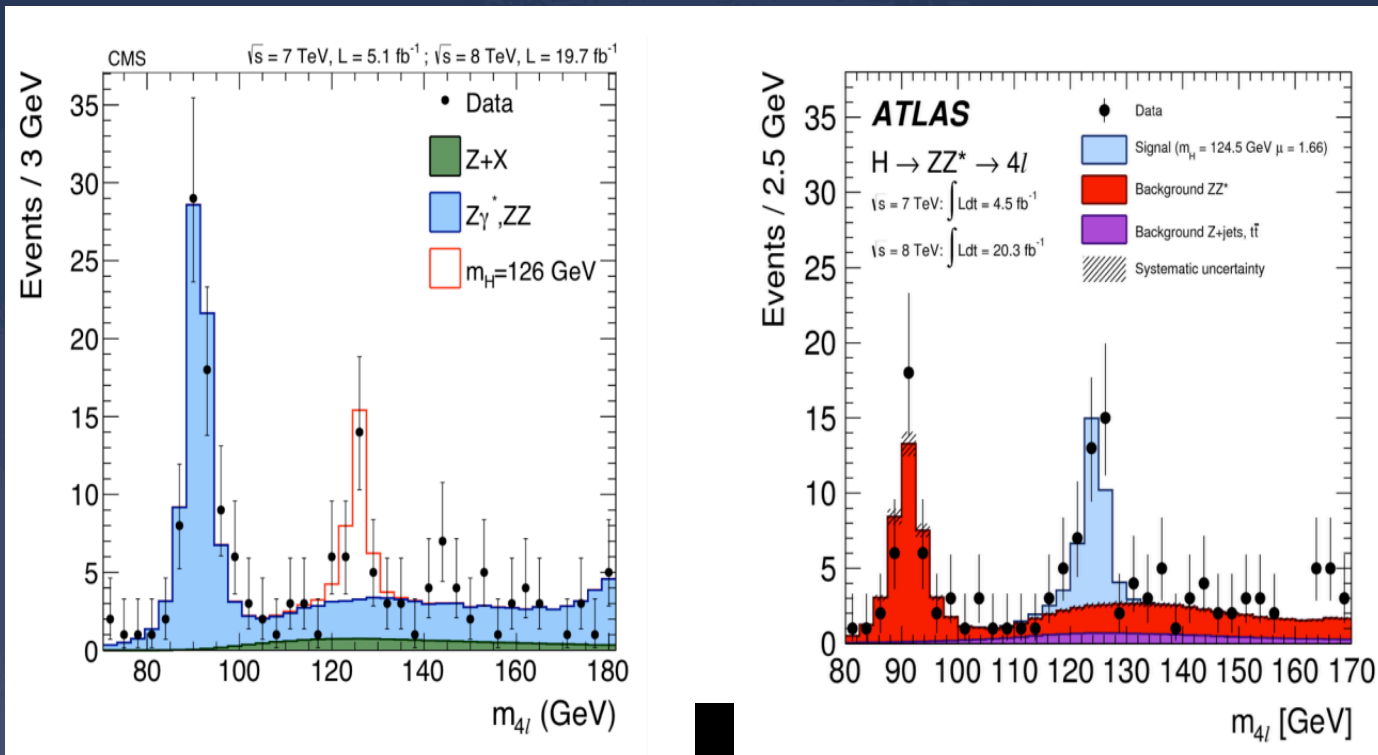
H1  $6\sigma$  pentaquarks  
hep-ex/0403017



CDF  $4\sigma$  dijet bump  
<http://arxiv.org/abs/1104.0699>

# And Resonances That Are!

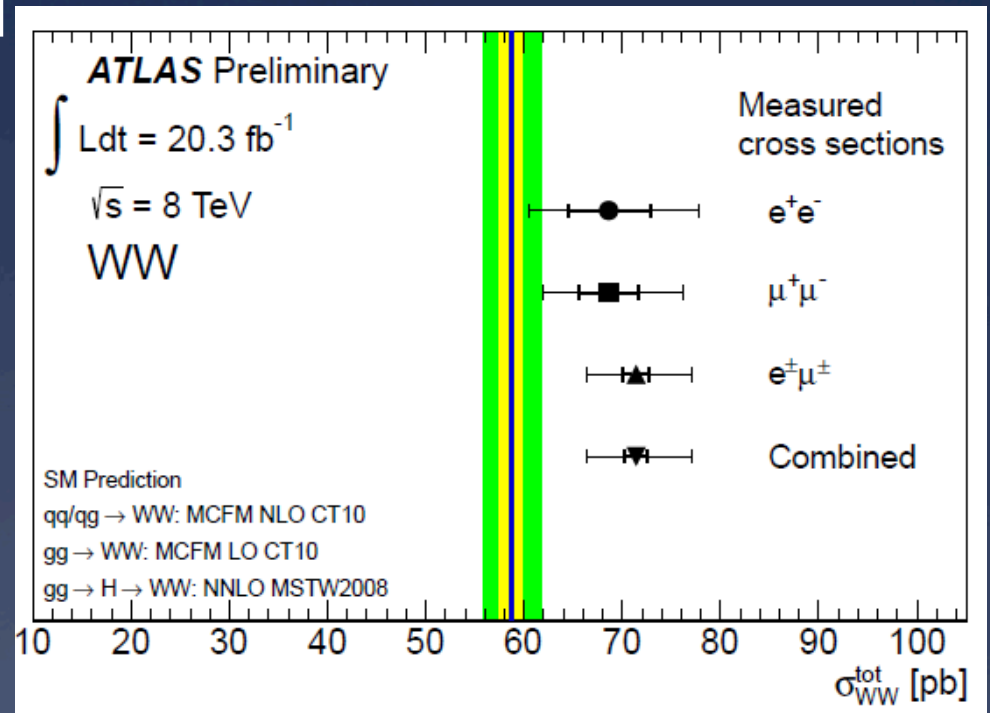
CMS, ATLAS, Higgs boson,  $> 5\sigma$



# A (Cautionary) Tale of an Excess

\* CMS and ATLAS measured a cross section  $\sim 20\%$  higher than the NLO SM prediction.

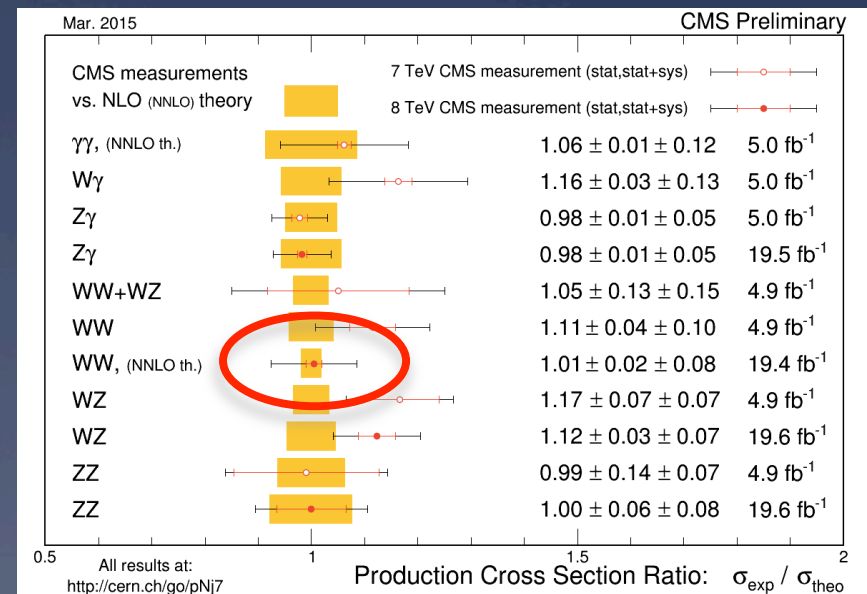
\* This exciting result triggered a frenzy of speculation:  
arXiv:1406.0848  
arXiv:1406.0858  
arXiv:1407.4912, etc.



# A (Cautionary) Tale of an Excess

\* My colleagues Jaiswal and Okui (arXiv:1407.4537), and Meade, Ramani, and Zeng (arXiv:1407.4537) noted that the jet veto cut  $p_T > p_{T,veto}$ , used by ATLAS and CMS to reduce background from top quark production, renders the NLO calculation inaccurate because terms like  $\log^n(M/p_{T,veto})$  now need to be included. Improved calculations brought better agreement.

\* Finally, a NNLO calculation of the WW cross section (Phys. Rev. Lett. 113, 212001 (2014)) is found to agree with the measurements:



# So How Sure is Sure Enough?

Our field uses a rigid  $5\sigma$  rule to claim a discovery. But, is this sensible?

Louis Lyons argues perhaps not:

<http://arxiv.org/pdf/1310.1284v1.pdf>.

The point is whether you accept a statistically significant hypothesis  $H_1$  depends on your prior opinion ( $\pi$ ) about that hypothesis:

$$\frac{p(H_1 | data)}{p(H_0 | data)} = \frac{p(data | H_1) \pi_1}{p(data | H_0) \pi_0}$$

# So How Sure is Sure Enough?

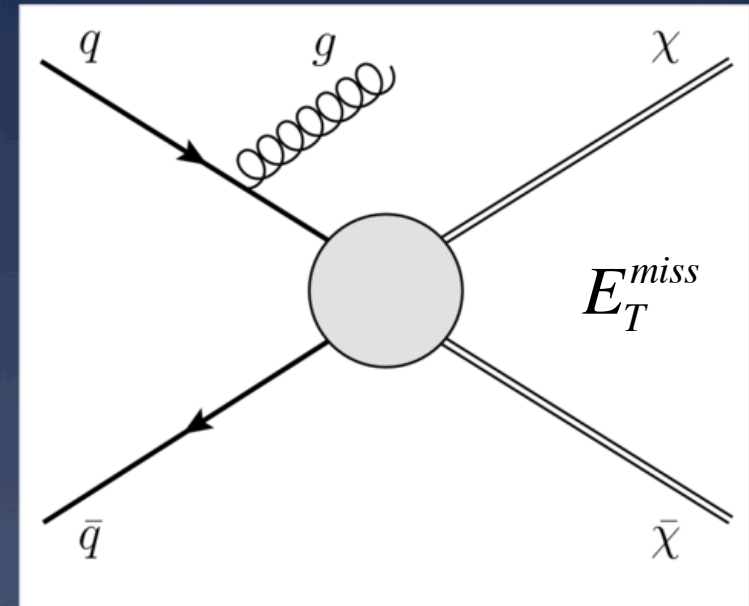
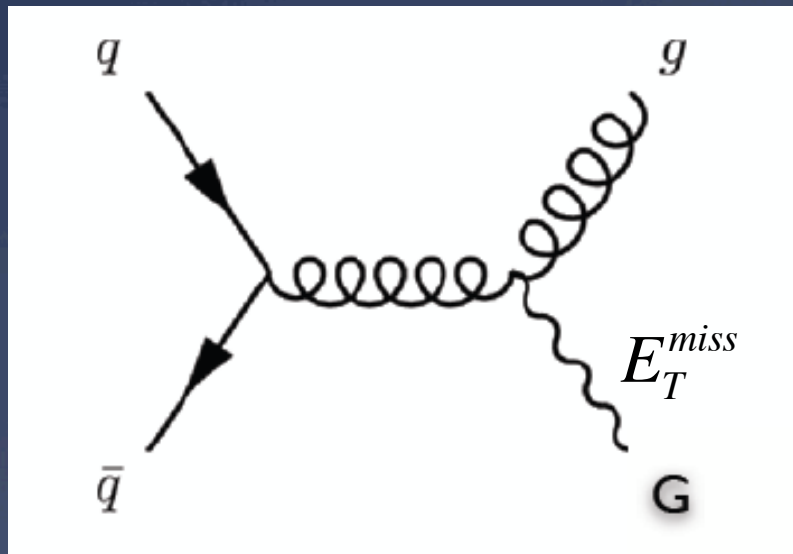
<http://arxiv.org/pdf/1310.1284v1.pdf>

Search	Degree of surprise	Impact	LEE	Systematics	Number of $\sigma$
Higgs search	Medium	Very high	Mass	Medium	5
Single top	No	Low	No	No	3
SUSY	Yes	Very high	Very large	Yes	7
$B_s$ oscillations	Medium/low	Medium	$\Delta m$	No	4
Neutrino oscillations	Medium	High	$\sin^2(2\theta), \Delta m^2$	No	4
$B_s \rightarrow \mu\mu$	No	Low/Medium	No	Medium	3
Pentaquark	Yes	High/very high	M, decay mode	Medium	7
$(g-2)_\mu$ anomaly	Yes	High	No	Yes	4
H spin $\neq 0$	Yes	High	No	Medium	5
$A^{\text{th}}$ generation $g, l, \nu$	Yes	High	M, mode	No	6
$v_\nu > c$	Enormous	Enormous	No	Yes	>8
Dark matter (direct)	Medium	High	Medium	Yes	5
Dark energy	Yes	Very high	Strength	Yes	5
Grav waves	No	High	Enormous	Yes	7

# The Mono X Program

or Clapping With One Hand!

Looking for “nothing”  
recoiling against  
something ( $X = \text{e.g., a jet}$ ).



# The Mono X Program

Example: 8 TeV ATLAS mono-jet search (1502.01518, 15 Sep 2015) imposed the requirements

$$p_{Tj} > 30 \text{ GeV}, |\eta_j| < 4.5$$

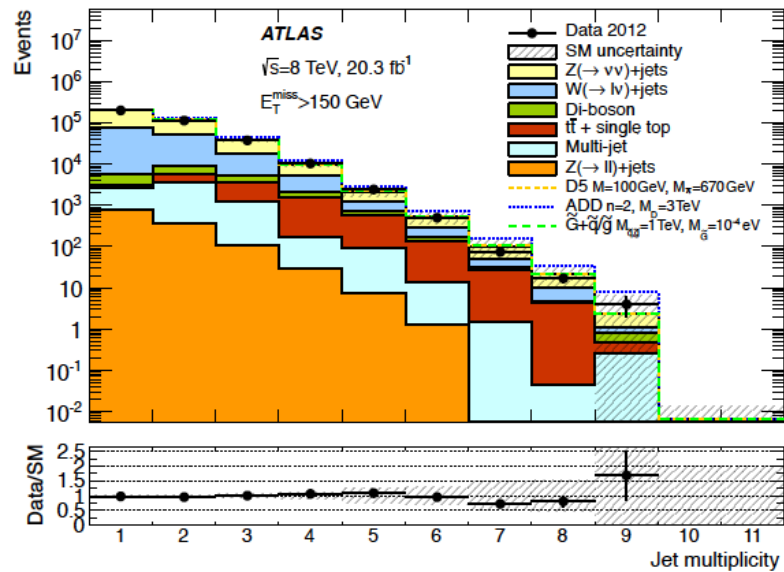
$$p_{T1} > 120 \text{ GeV}, |\eta_{j1}| < 2.0, p_{T1} / E_T^{mis} > 0.5$$

$$\Delta\phi(\text{jet}, \vec{p}_T^{miss}) > 1.0$$

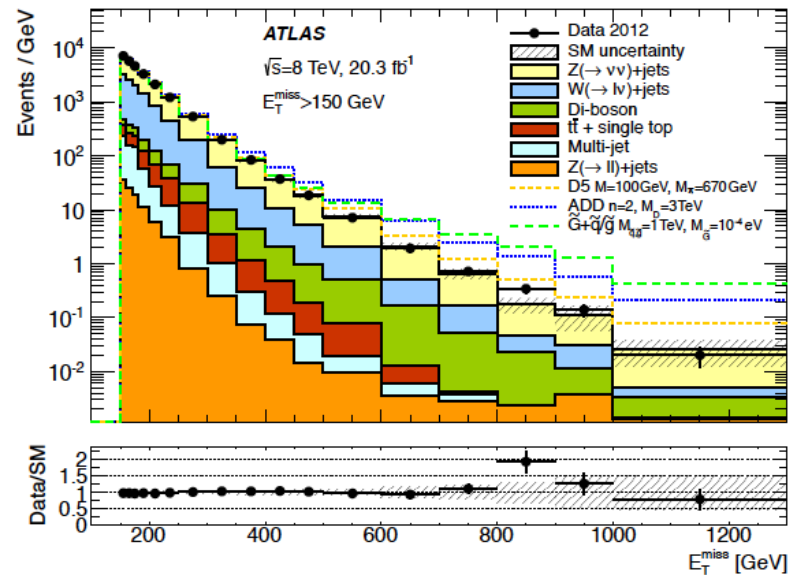
to select mono-jet like events.

All backgrounds (single-boson, di-boson, top quark) are simulated, except for multi-jets and non-collision events, which are estimated using data-driven methods.

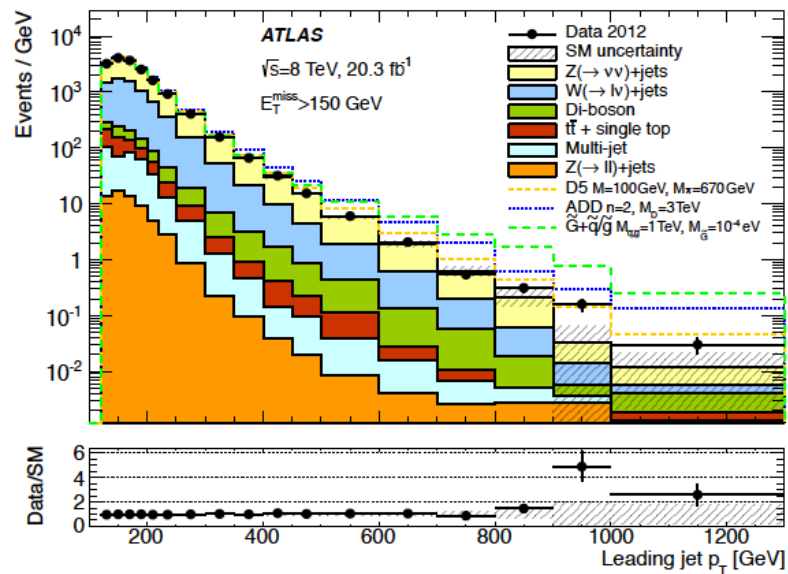




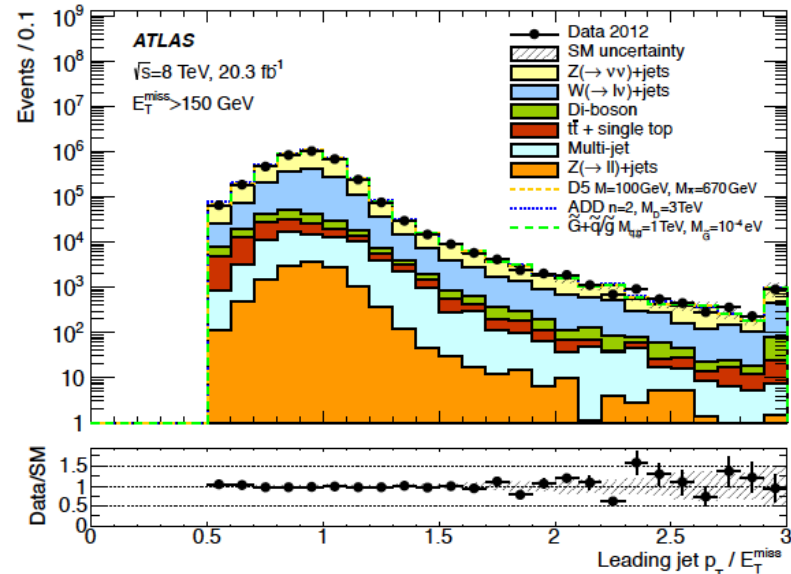
(a)



(b)



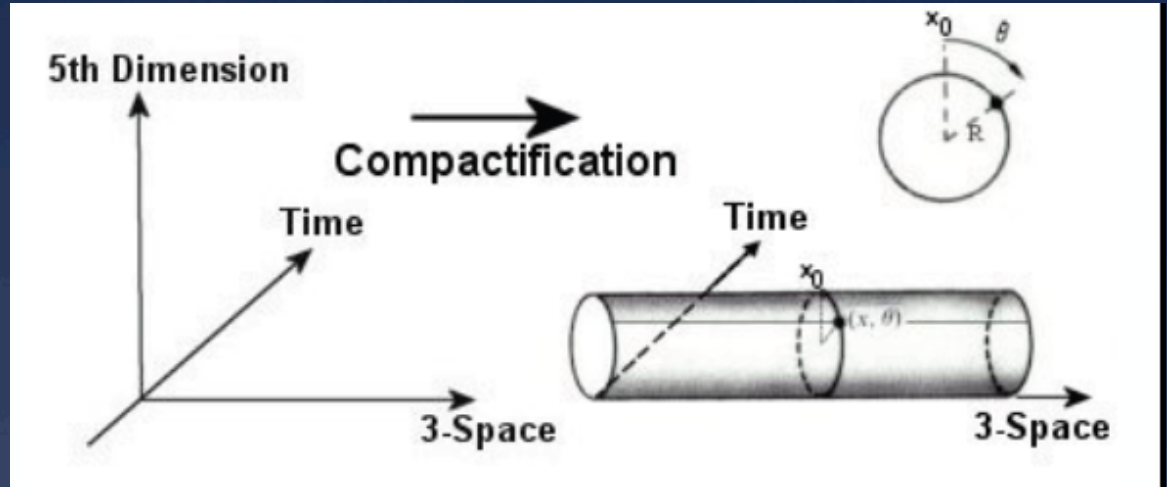
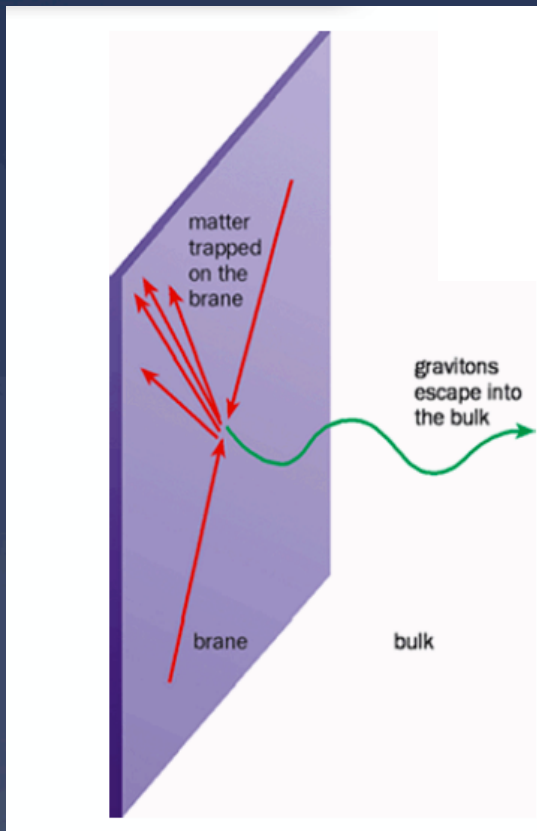
(c)



(d)

# The Mono X Program

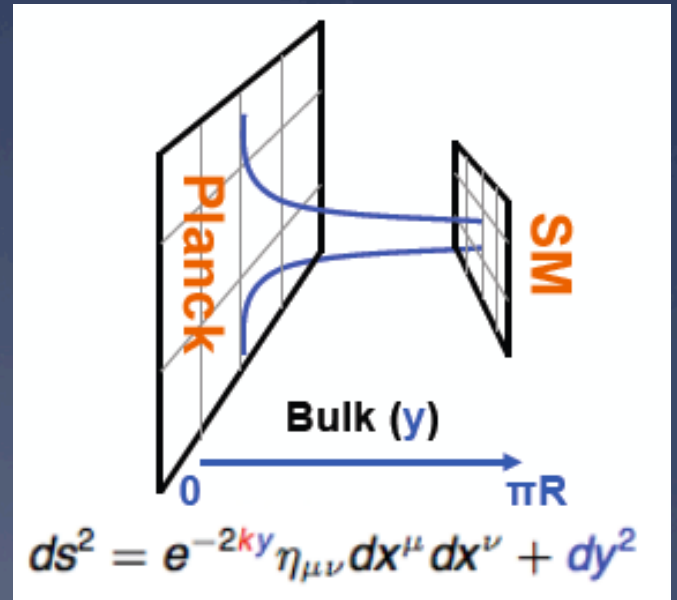
Arkani-Hamed  
Dimopoulos  
Dvali



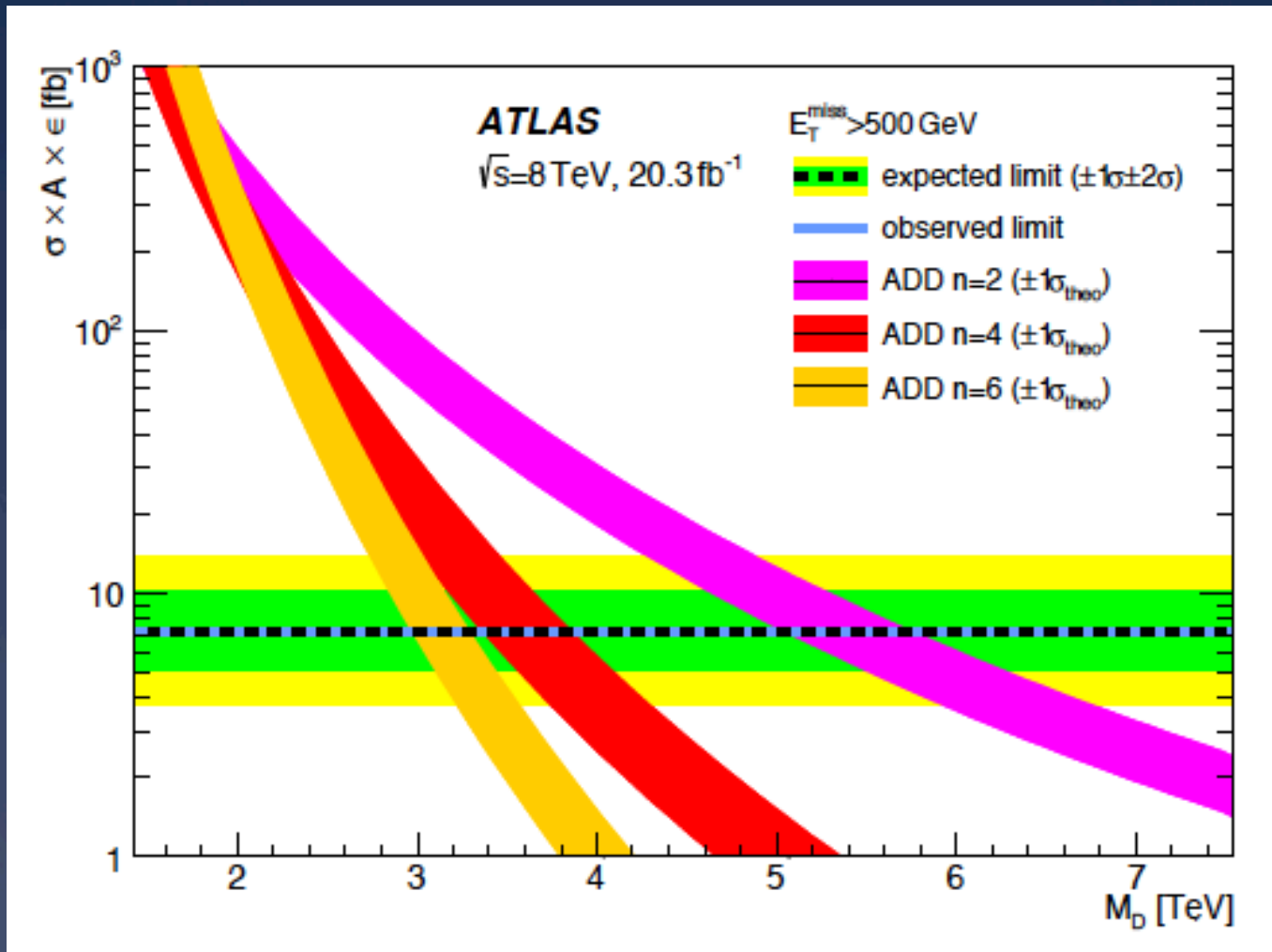
Randall  
Sundrum

$$\text{ADD } M_{Pl} \sim M_{EWK}^{1+n/2} R^{n/2}$$

$$\text{RS } M_{Pl} \sim \Lambda e^{k\pi R}$$



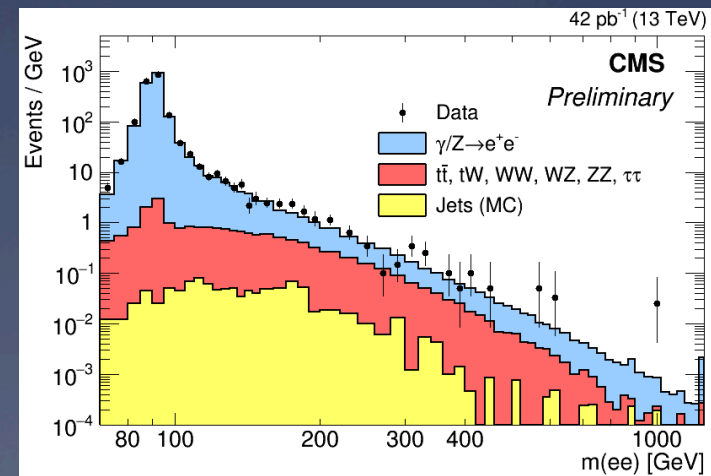
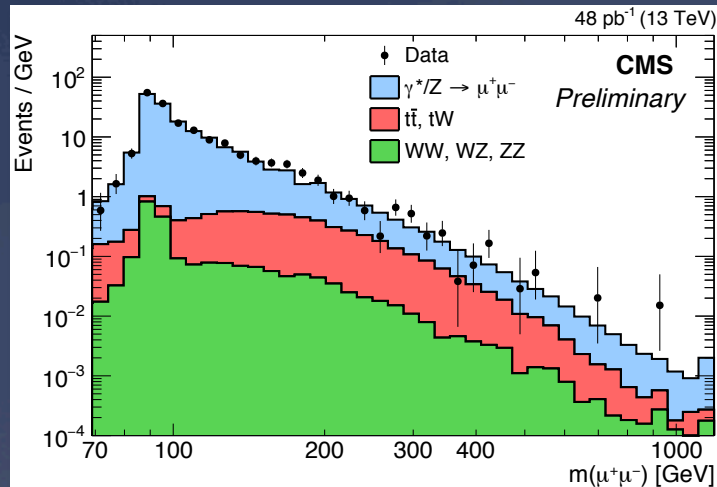
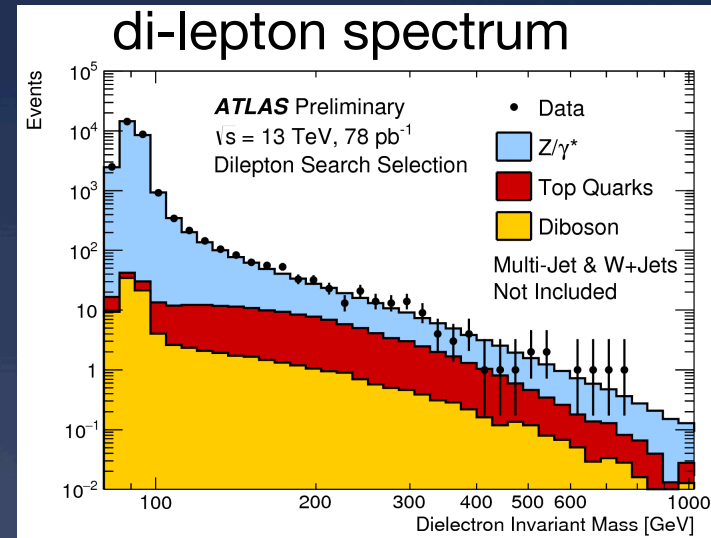
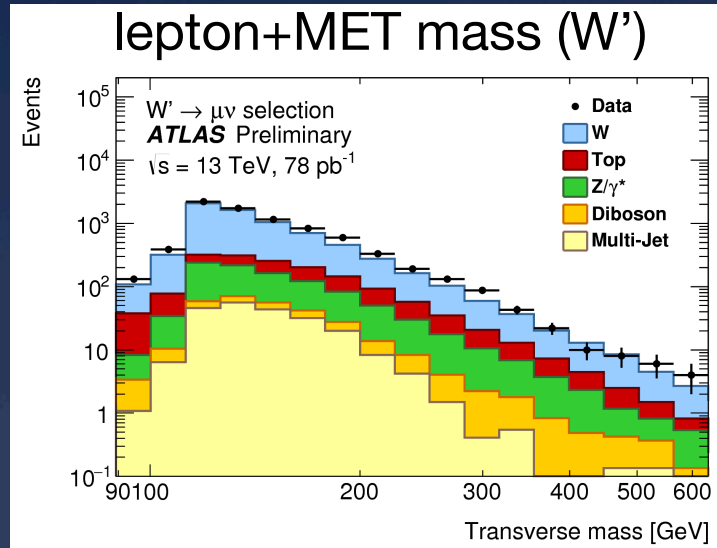
# The Mono X Program



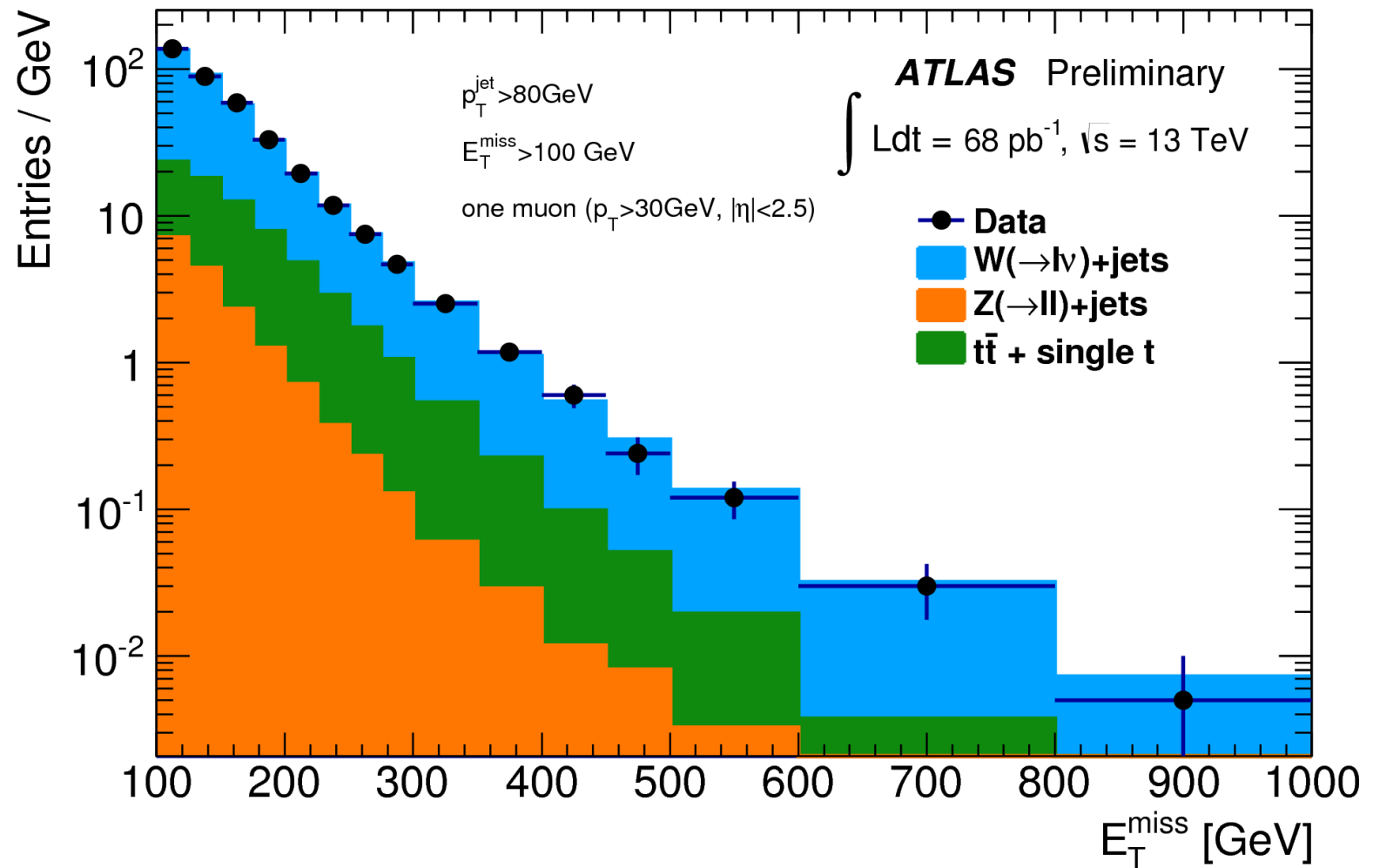
# Outline

- \* Introduction
- \* LHC Run I – A Random Stroll
- \* **LHC Run II – Early Results**
- \* Opportunities and Challenges

# Early Commissioning Results



# ATLAS@13 TeV: Breathtaking!



# Resonance Searches – Round 2

CMS search for a narrow resonance in dijet mass spectrum

$$p_T > 30 \text{ GeV}$$

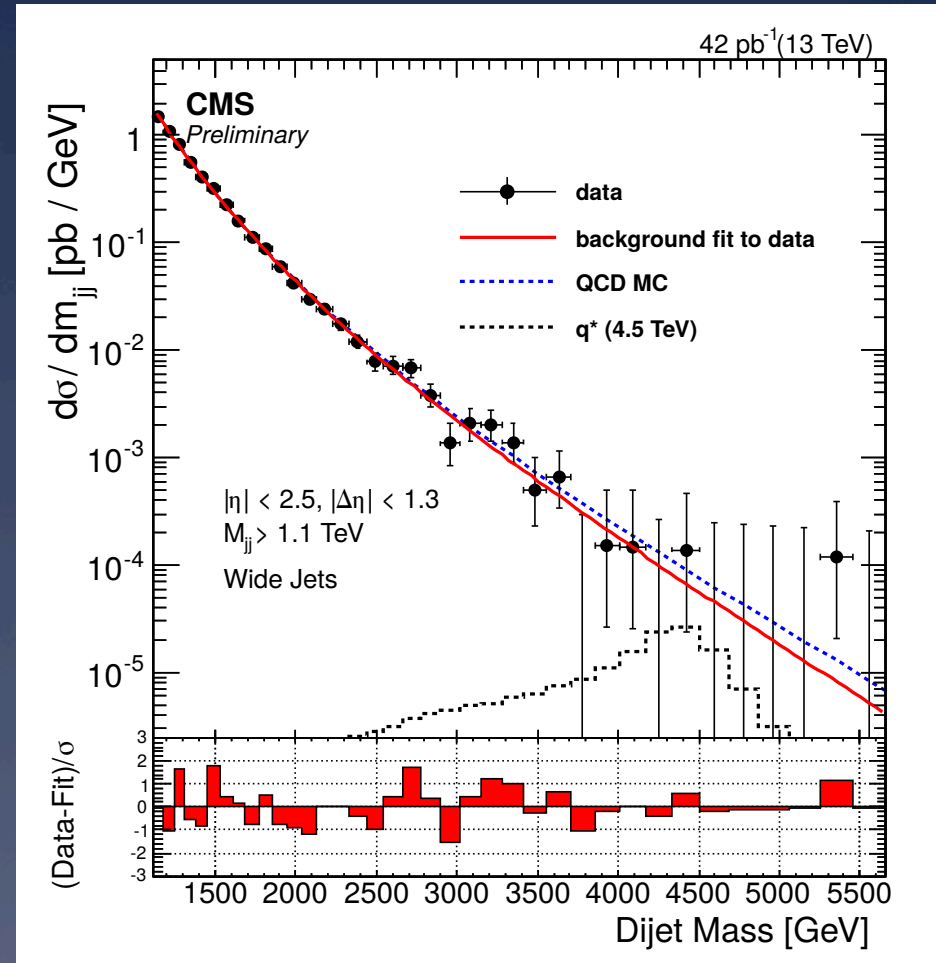
$$|\eta_j| < 2.5$$

$$|\Delta\eta_{jj}| > 1.3$$

$$\Delta R = \sqrt{\Delta\eta^2 + \Delta\phi^2} < 1.1$$

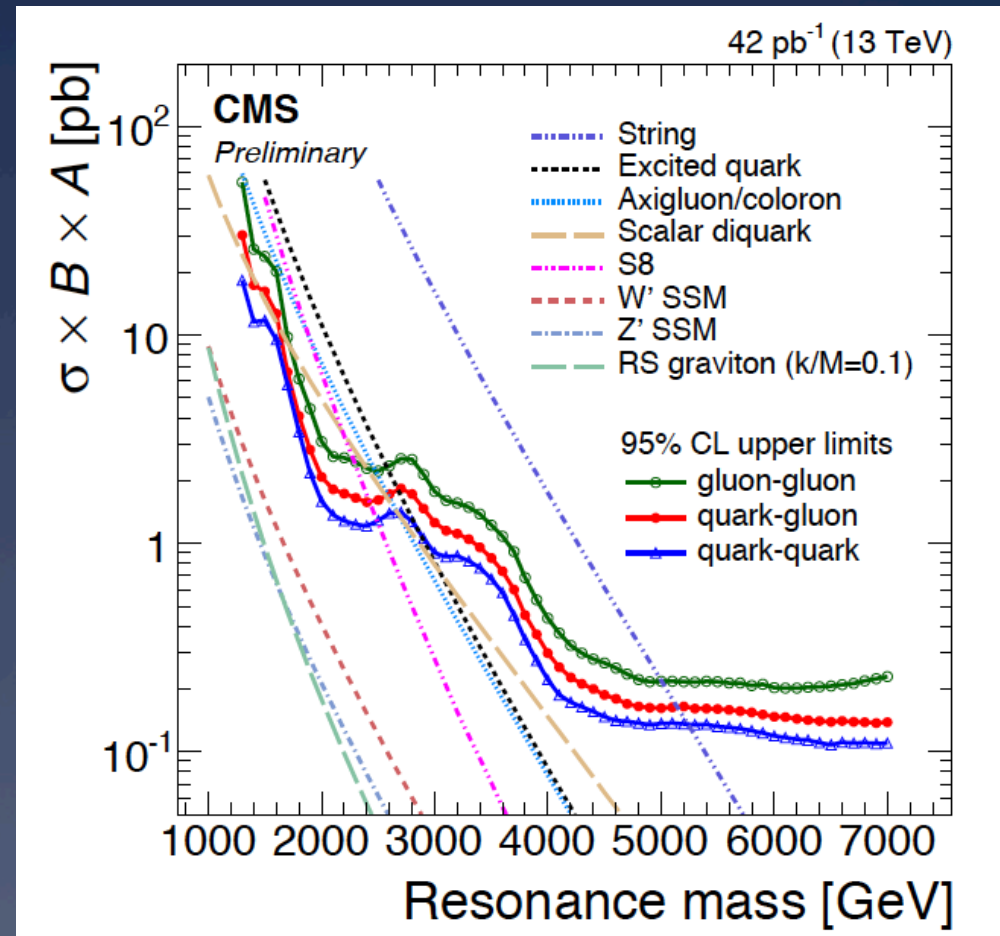
Background modeled with

$$\frac{d\sigma}{m_{jj}} = p_1(1-x)^{p_2} x^{-p_3}$$



# Resonance Searches – Round 2

- \* As expected, this search can be used to set limits on many models:
- \* But, yet again, the Standard Model reigns supreme!





# A 2.9 TeV Di-Electron Event

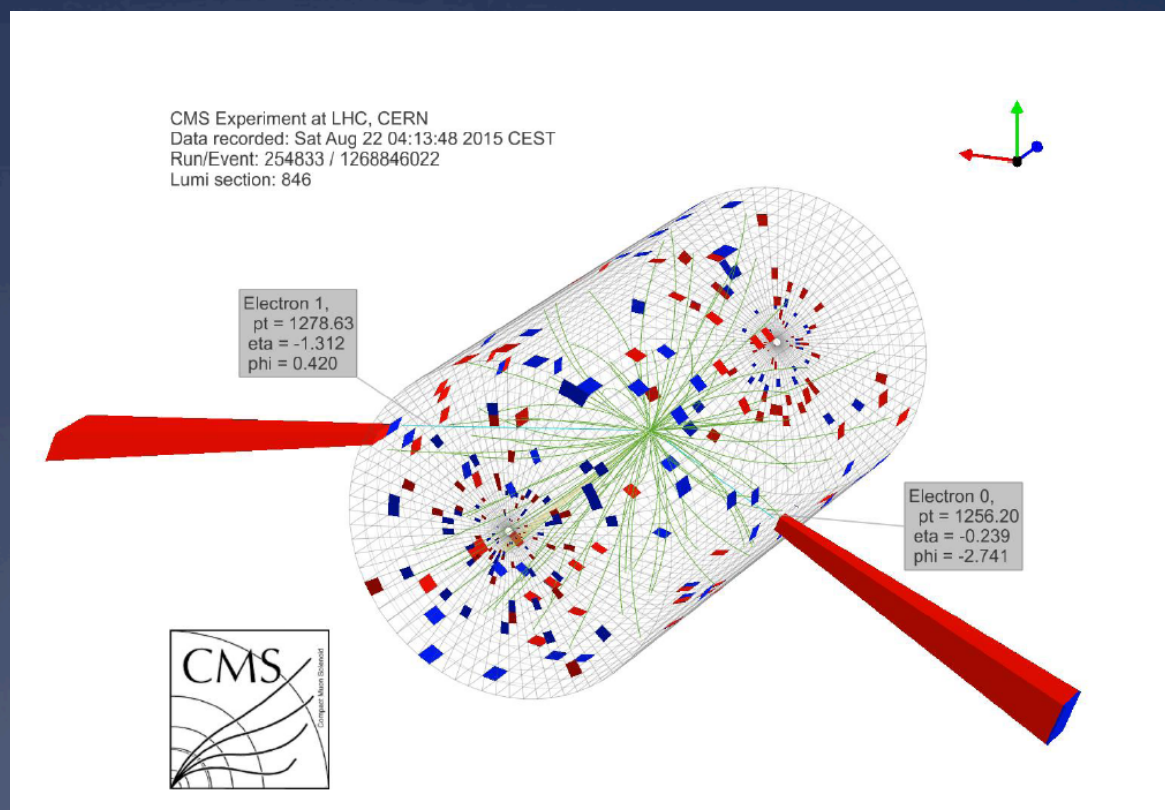
$$p_{T1} = 1.28 \text{ TeV}, \eta_1 = -1.31, \phi_1 = 0.42$$

$$p_{T2} = 1.26 \text{ TeV}$$

$$\eta_2 = -1.24,$$

$$\phi_2 = -2.74$$

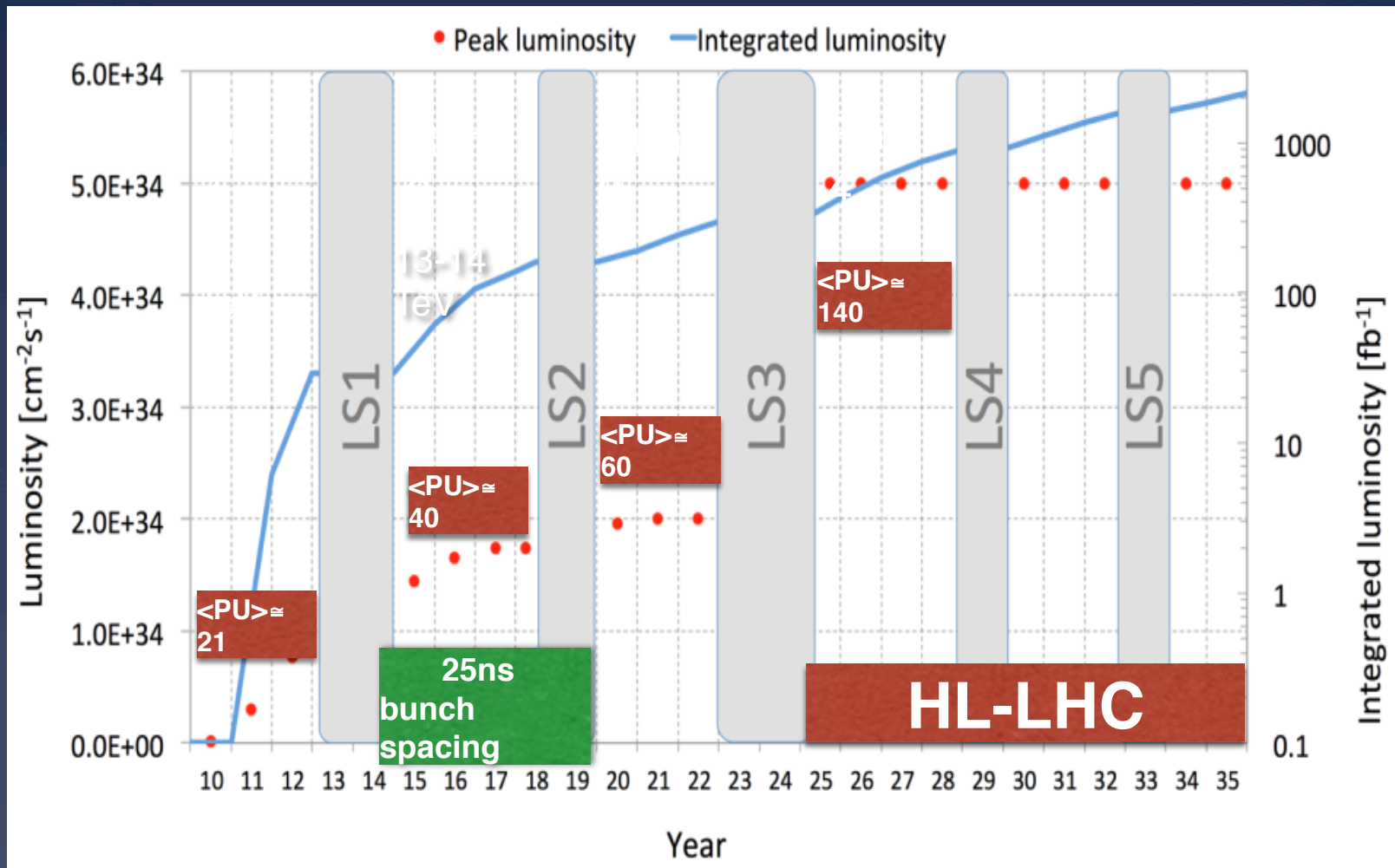
Expected  
background  
above 2 TeV  
0.007 events



# Outline

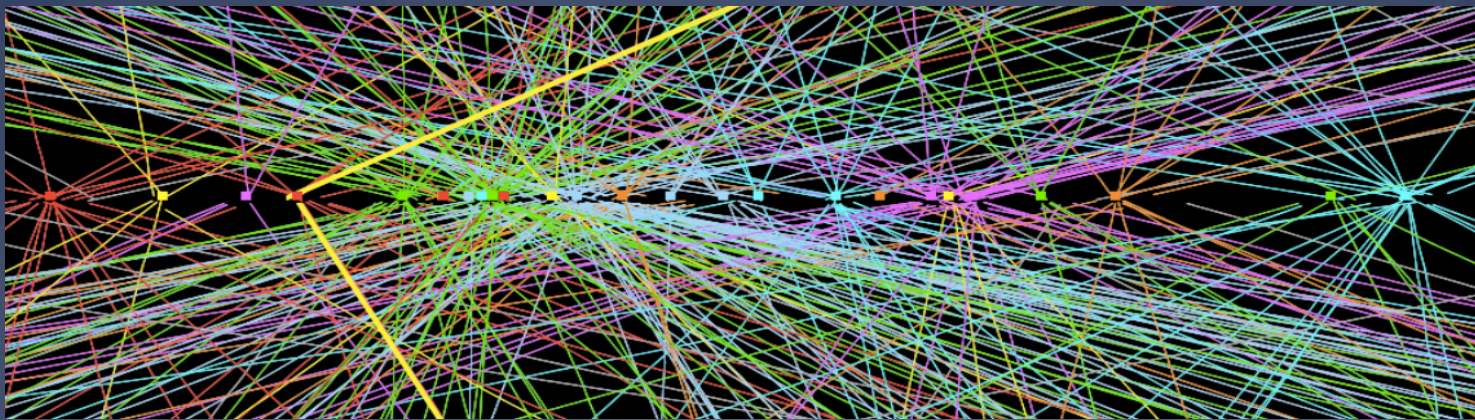
- \* Introduction
- \* LHC Run I – A Random Stroll
- \* LHC Run II – Early Results
- \* **Opportunities and Challenges**

# Opportunities & Challenges



# Opportunities & Challenges

- \* The upgraded LHC and its amazing detectors provide the best chance we have to move the field significantly forward.
- \* The opportunities for young physicists to make lasting contributions are enormous. They are so precisely because of the many challenges we must face!
- \* Take, for example, pileup:



# Opportunities & Challenges

At  $L \sim 1.3 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ , find  $\sim 140$  vertices spread over 5 cm!

$$\mu = \frac{\sigma_{tot} L}{n_b f}$$

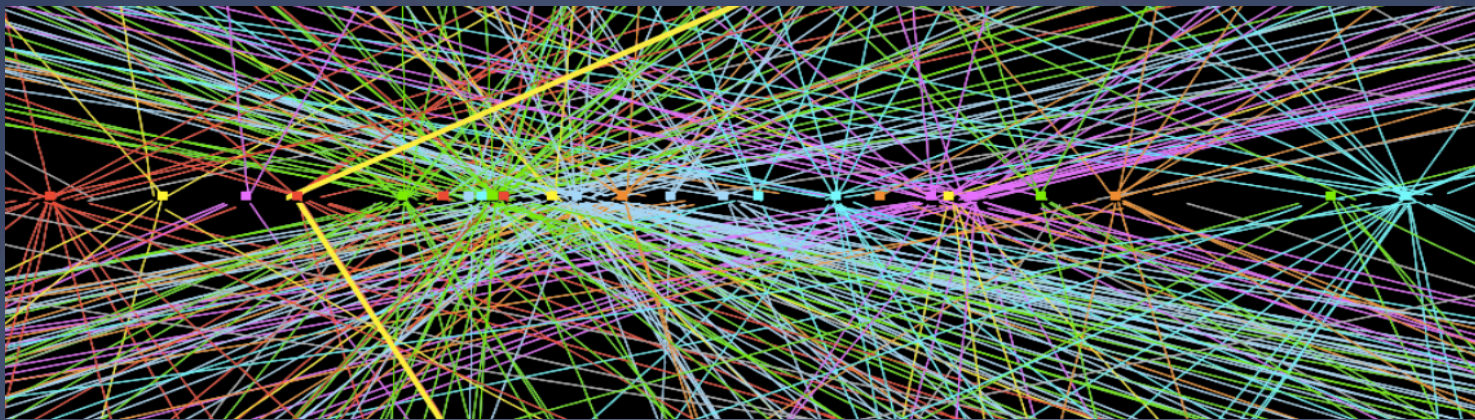
$\mu$  average number of events/bunch crossing

$\sigma_{tot}$  total inelastic cross section

$L$  luminosity

$n_b$  number of bunches

$f$  bunch collision frequency



# Opportunities & Challenges

## A Few Experimental Challenges

- \* Handle data storage at the rate of 6 gigabytes/second
- \* Trigger efficiently, even at low thresholds
- \* Mitigate the effects of pileup
- \* Efficient identification of boosted W, Z, Higgs, top
- \* Efficient identification of b jets and taus
- \* High fidelity, yet fast, detector simulation
- \* Automatic search for deviations from SM

## A Couple of Theoretical Physics Challenges (my wish list!)

- \* Fully automatic high precision predictions for hundreds of SM exclusive final states.
- \* Precise calculation of multijet cross sections, finally!

# Opportunities & Challenges

- \* Where does our field want to be in 2050?
- \* We want to be in possession of the New Standard Model.
- \* However, for that to happen, two things are necessary:
  - \* We must discover new physics, and
  - \* we must significantly narrow the divide between theorists and experimentalists: *We Need To Talk More*.
- \* But, for the latter to happen, a major cultural shift will be necessary.

# Final Remarks

The Standard Model – a complicated Heath Robinson edifice ingeniously put together by brilliant architects and builders – is to be celebrated as a truly magnificent achievement.

But the **NSM** will be even more extraordinary.



William Heath  
Robinson  
1872 – 1944

