

Higgs Boson Phenomenology Lecture II

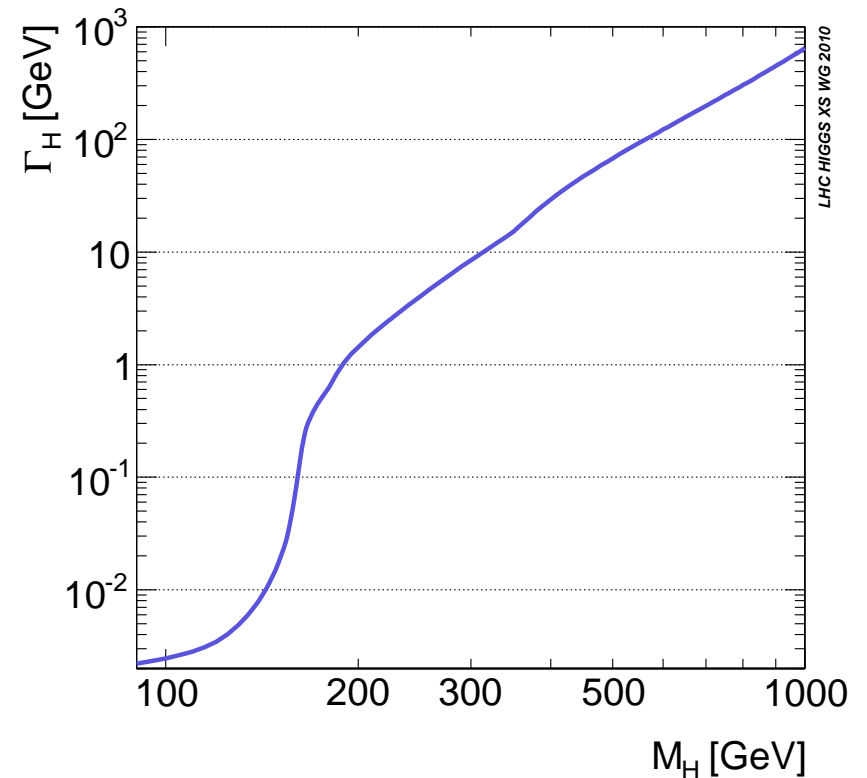
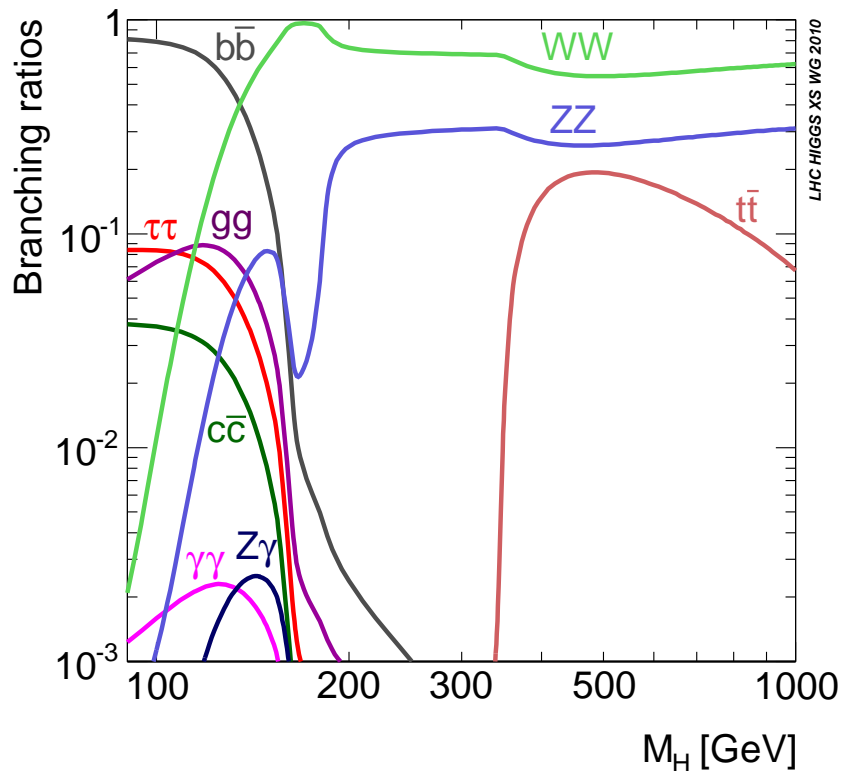
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Outline of Lecture II

- SM Higgs boson physics depends just on M_H : M_H highly constrained!
- **Quantum effects** leading players in constraining M_H :
 - branching ratios (see end of last lecture): important corrections;
 - EW precision fits: M_H only unknown.
- SM Higgs so constrained that it **can point to scale of new physics**.
- **Beyond SM Higgs**: MSSM (useful example of an extended Higgs sector) and more.
- Need data! and need to understand them

SM Higgs boson decay branching ratios and width



Observe difference between light and heavy Higgs

These curves include: tree level + QCD and EW loop corrections.

Tree level decays: $H \rightarrow f\bar{f}$ and $H \rightarrow VV$

At lowest order:

$$\begin{aligned}\Gamma(H \rightarrow f\bar{f}) &= \frac{G_F M_H}{4\sqrt{2}\pi} N_{cf} m_f^2 \beta_f^3 \\ \Gamma(H \rightarrow VV) &= \frac{G_F M_H^3}{16\sqrt{2}\pi} \delta_V \left(1 - \tau_V + \frac{3}{4}\tau_V^2\right) \beta_V\end{aligned}$$

$$(\beta_i = \sqrt{1 - \tau_i}, \tau_i = 4m_i^2/M_H^2, \delta_{W,Z} = 2, 1, (N_c)_{l,q} = 1, 3)$$

Ex.1: Higher order corrections to $H \rightarrow q\bar{q}$

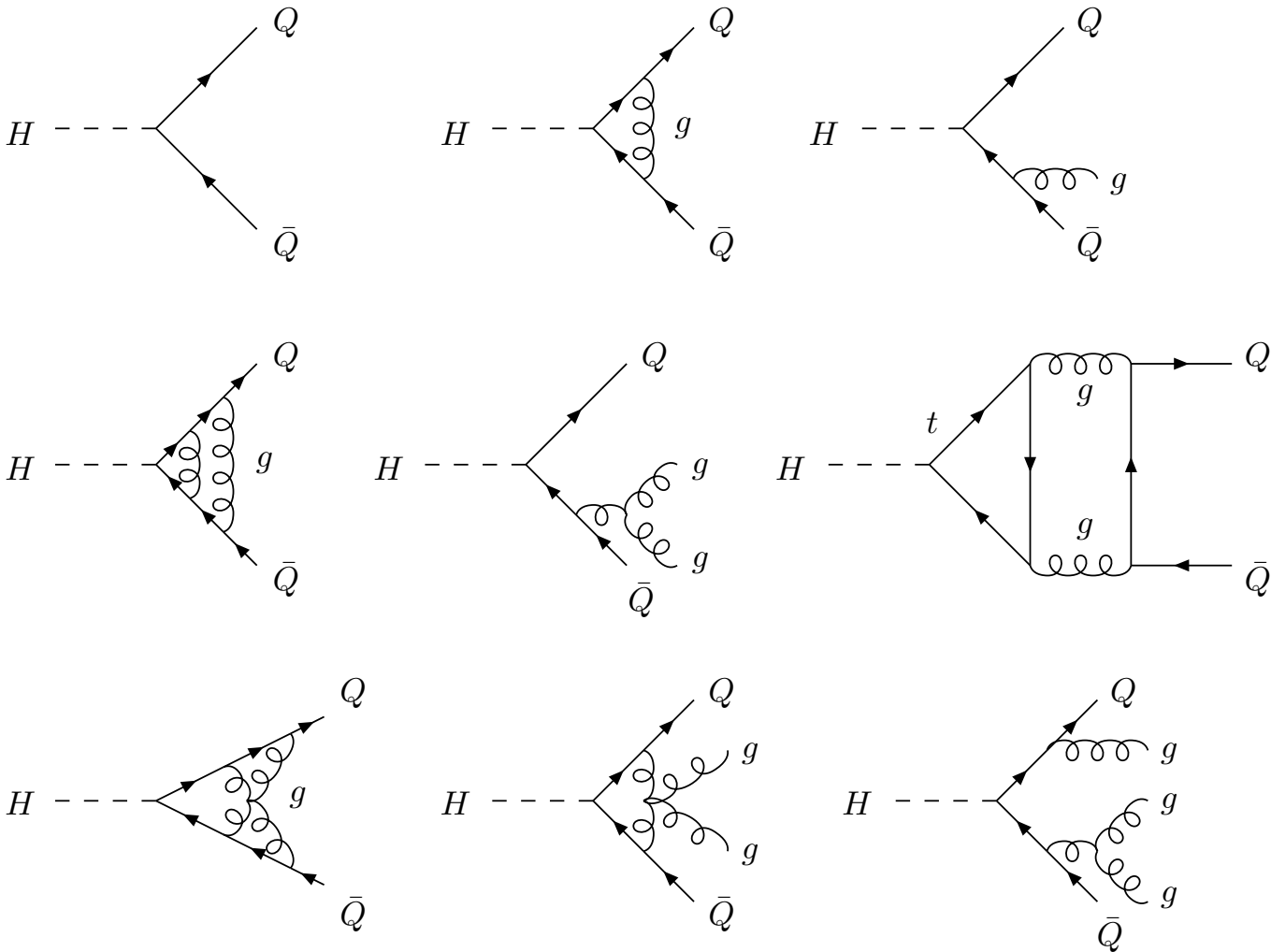
QCD corrections dominant:

$$\Gamma(H \rightarrow q\bar{q})_{\text{QCD}} = \frac{3G_F M_H}{4\sqrt{2}\pi} \bar{m}_q^2(M_H) \beta_q^3 [\Delta_{\text{QCD}} + \Delta_t]$$

$$\Delta_{\text{QCD}} = 1 + 5.67 \frac{\alpha_s(M_H)}{\pi} + (35.94 - 1.36N_F) \left(\frac{\alpha_s(M_H)}{\pi}\right)^2 + \dots$$

$$\Delta_t = \left(\frac{\alpha_s(M_H)}{\pi}\right)^2 \left[1.57 - \frac{2}{3} \ln \frac{M_H^2}{m_t^2} + \frac{1}{9} \ln^2 \frac{\bar{m}_q^2(M_H)}{M_H^2}\right] + \dots$$

Consist of both virtual and real corrections, e.g.:



- Large Logs absorbed into \overline{MS} quark mass

$$\text{Leading Order: } \bar{m}_Q(\mu) = \bar{m}_Q(m_Q) \left(\frac{\alpha_s(\mu)}{\alpha_s(m_Q)} \right)^{\frac{2b_0}{\gamma_0}}$$

$$\text{Higher order: } \bar{m}_Q(\mu) = \bar{m}_Q(m_Q) \frac{f(\alpha_s(\mu)/\pi)}{f(\alpha_s(m_Q)/\pi)}$$

where (from renormalization group equation)

$$f(x) = \left(\frac{25}{6} x \right)^{\frac{12}{25}} [1 + 1.014x + \dots] \quad \text{for } m_c < \mu < m_b$$

$$f(x) = \left(\frac{23}{6} x \right)^{\frac{12}{23}} [1 + 1.175x + \dots] \quad \text{for } m_b < \mu < m_t$$

$$f(x) = \left(\frac{7}{2} x \right)^{\frac{4}{7}} [1 + 1.398x + \dots] \quad \text{for } \mu > m_t$$

- Large corrections, when $M_H \gg m_Q$

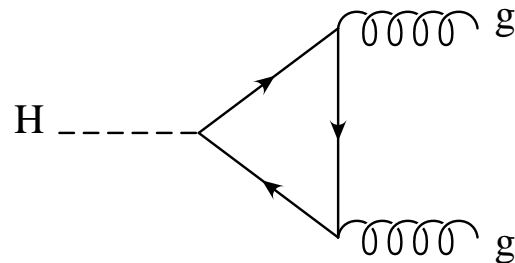
$$m_b(m_b) \simeq 4.2 \text{ GeV} \longrightarrow \bar{m}_b(M_h \simeq 100 \text{ GeV}) \simeq 3 \text{ GeV}$$

Branching ratio smaller by almost a factor 2.

- Main uncertainties: $\alpha_s(M_Z)$, pole masses: $m_c(m_c)$, $m_b(m_b)$.

Ex. 2: Higher order corrections $\Gamma(H \rightarrow gg)$

Start from tree level:



$$\Gamma(H \rightarrow gg) = \frac{G_F \alpha_s^2 M_H^3}{36\sqrt{2}\pi^3} \left| \sum_q A_q^H(\tau_q) \right|$$

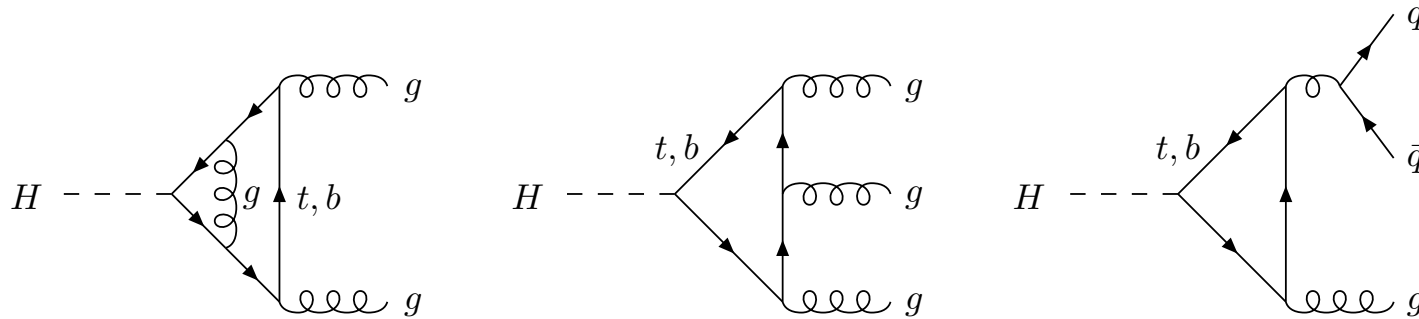
where $\tau_q = 4m_q^2/M_H^2$ and

$$A_q^H(\tau) = \frac{3}{2}\tau [1 + (1 - \tau)f(\tau)]$$

$$f(\tau) = \begin{cases} \arcsin^2 \frac{1}{\sqrt{\tau}} & \tau \geq 1 \\ -\frac{1}{4} \left[\ln \frac{1+\sqrt{1-\tau}}{1-\sqrt{1-\tau}} - i\pi \right]^2 & \tau < 1 \end{cases}$$

Main contribution from top quark \rightarrow optimal situation to use **Low Energy Theorems** to add higher order corrections.

QCD corrections dominant:



Difficult task since decay is already a loop effect.

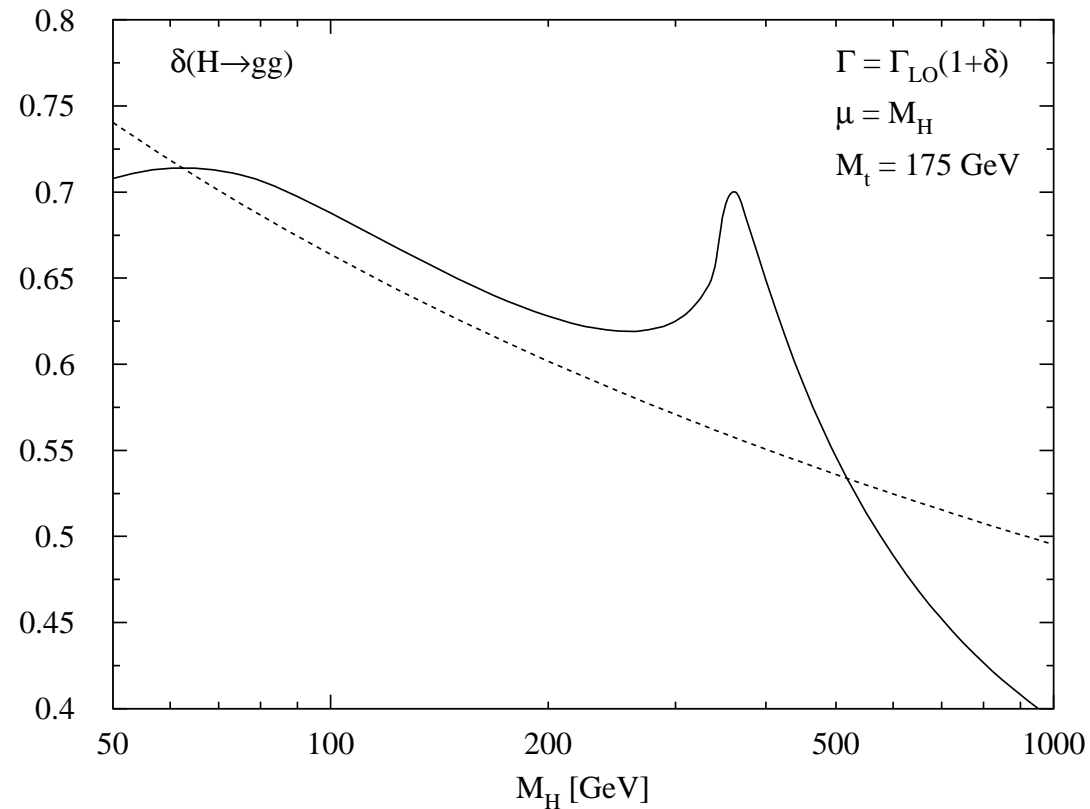
However, full massive calculation of $\Gamma(H \rightarrow gg(q), q\bar{q}g)$ agrees with $m_t \gg M_H$ result at 10%

$$\Gamma(H \rightarrow gg(q), q\bar{q}g) = \Gamma_{LO}(\alpha_s^{(N_L)}(M_H)) \left[1 + E^{(N_L)} \frac{\alpha_s^{(N_L)}}{\pi} \right]$$

$$E^{(N_L)} \xrightarrow{M_H^2 \ll 4m_q^2} \frac{95}{4} - \frac{7}{6}N_L$$

Dominant soft/collinear radiation do not resolve the Higgs boson coupling to gluons \rightarrow QCD corrections are just a (big) rescaling factor

NLO QCD corrections almost 60 – 70% of LO result in the low mass region:



solid line \longrightarrow full massive NLO calculation

dashed line \longrightarrow heavy top limit ($M_H^2 \ll 4m_t^2$)

NNLO corrections calculated in the heavy top limit: add 20%

\longrightarrow perturbative stabilization. Residual theoretical uncertainty $\simeq 10\%$.

Low-energy theorems, in a nutshell.

- Observing that:

In the $p_H \rightarrow 0$ limit: the interactions of a Higgs boson with the SM particles arise by substituting

$$M_i \longrightarrow M_i \left(1 + \frac{H}{v} \right) \quad (i = f, W, Z)$$

In practice: Higgs taken on shell ($p_H^2 = M_H^2$), and limit $p_H \rightarrow 0$ is limit of small Higgs masses (e.g.: $M_H^2 \ll 4m_t^2$).

- Then

$$\lim_{p_H \rightarrow 0} \mathcal{A}(X \rightarrow Y + H) = \frac{1}{v} \sum_i M_i \frac{\partial}{\partial M_i} \mathcal{A}(X \rightarrow Y)$$

very convenient!

- Equivalent to an **Effective Theory** described by:

$$\mathcal{L}_{eff} = \frac{\alpha_s}{12\pi} G^{a\mu\nu} G_{\mu\nu}^a \frac{H}{v} (1 + O(\alpha_s))$$

including higher order QCD corrections.

For completeness:

$$\Gamma(H \rightarrow \gamma\gamma) = \frac{G_F \alpha^2 M_H^3}{128 \sqrt{2} \pi^3} \left| \sum_f N_{cf} e_f^2 A_f^H(\tau_f) + A_W^H(\tau_W) \right|^2$$

where $f(\tau)$ as in $H \rightarrow gg$:

$$\begin{aligned} A_f^H &= 2\tau [1 + (1 - \tau)f(\tau)] \\ A_W^H(\tau) &= -[2 + 3\tau + 3\tau(2 - \tau)f(\tau)] \end{aligned}$$

$$\Gamma(H \rightarrow Z\gamma) = \frac{G_F^2 M_W^2 \alpha M_H^3}{64 \pi^4} \left(1 - \frac{M_Z^2}{M_H^2}\right)^3 \left| \sum_f A_f^H(\tau_f, \lambda_f) + A_W^H(\tau_W, \lambda_W) \right|^2$$

where the form factors $A_f^H(\tau, \lambda)$ and $A_W^H(\tau, \lambda)$ can be found in the literature (see, e.g., M. Spira, hep-ph/9705337).

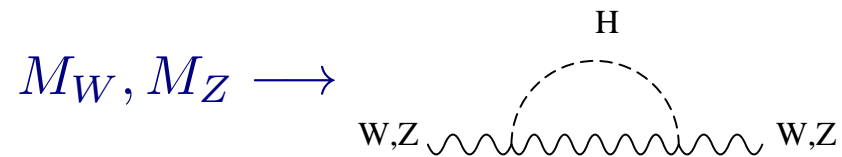
For both decays, both QCD and EW corrections are very small ($\simeq 1 - 3\%$).

EW precision fits: perturbatively calculate observables in terms of few parameters:

$$M_Z, G_F, \alpha(M_Z), M_W, m_f, (\alpha_s(M_Z))$$

extracted from experiments with high accuracy. Only SM unknown: M_H .

- SM needs **Higgs boson** to **cancel infinities**, e.g.



- Finite **logarithmic contributions** survive, e.g. radiative corrections to $\rho = M_W^2 / (M_Z^2 \cos^2 \theta_W)$:

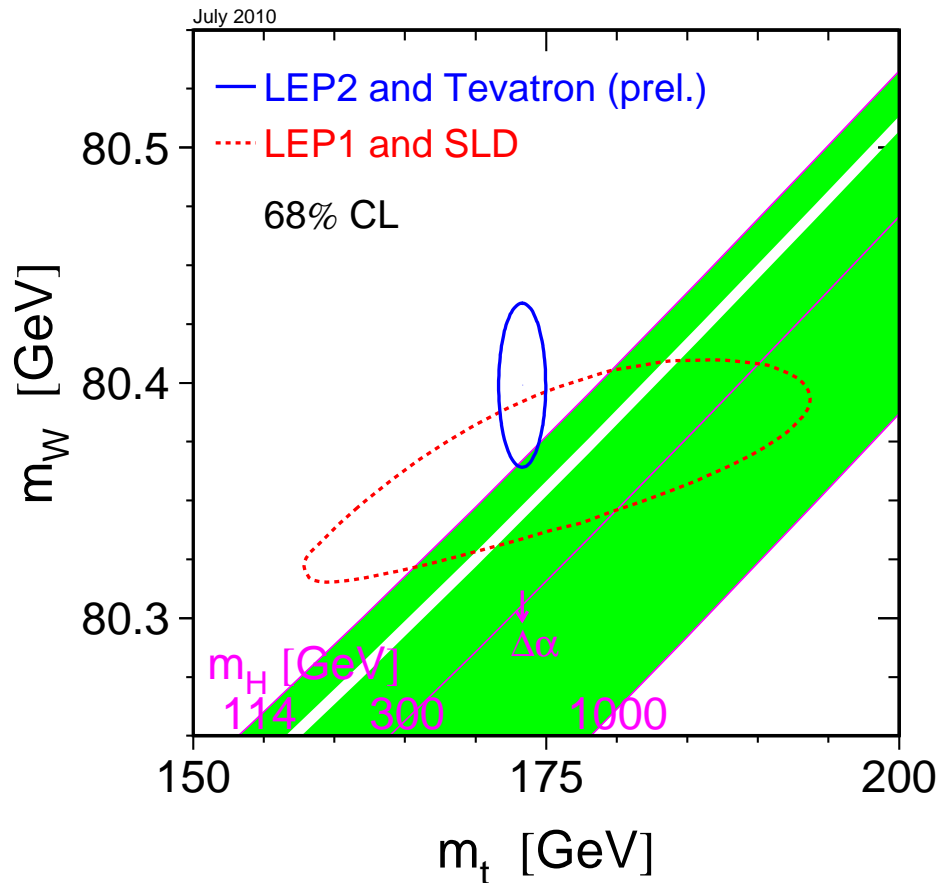
$$\rho = 1 - \frac{11g^2}{96\pi^2 \tan^2 \theta_W} \ln \left(\frac{M_H}{M_W} \right)$$

Main effects in **oblique radiative corrections** (S,T-parameters)

- Same constraints apply to any model of new physics.

SM Higgs-boson mass range: constrained by EW precision fits

Increasing precision will continue to provide an invaluable tool to test the consistency of the SM and its extensions.



$$m_W = 80.399 \pm 0.023 \text{ GeV}$$

$$m_t = 173.3 \pm 1.1 \text{ GeV}$$

↓

$$M_H = 89_{-26}^{+35} \text{ GeV}$$

$$M_H < 158 (185) \text{ GeV}$$

plus exclusion limits (95% c.l.):

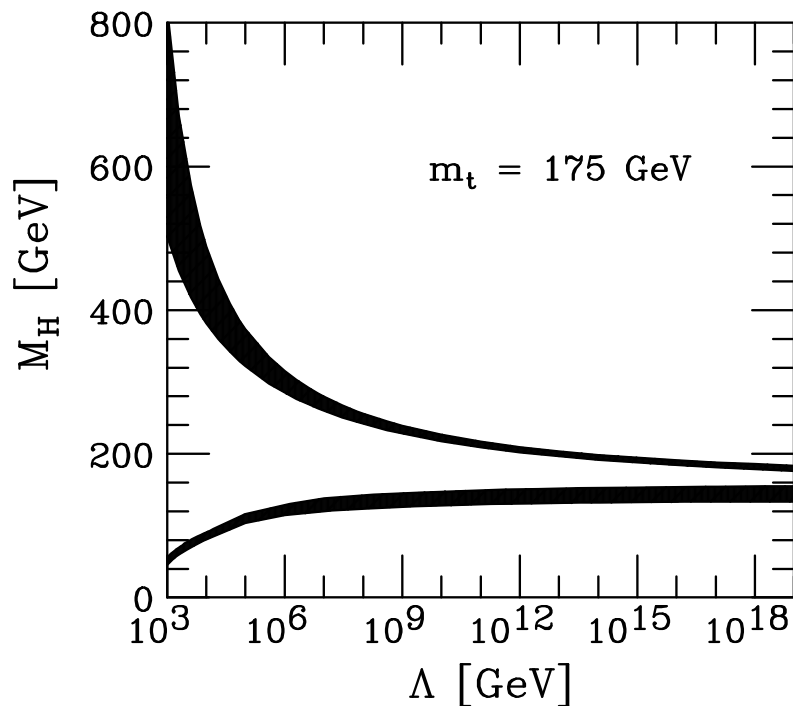
$$M_H > 114.4 \text{ GeV (LEP)}$$

$$M_H \neq 158 - 175 \text{ GeV (Tevatron)}$$

focus is now on exclusion limits and discovery!

Other theoretical constraints on M_H in the Standard Model

SM as an effective theory valid up to a scale Λ . The Higgs sector of the SM actually contains two unknowns: M_H and Λ .



Bounds given by:

- unitarity
- triviality
- vacuum stability
- fine tuning

$M_H^2 = 2\lambda v^2$ → M_H determines the weak/strong coupling behavior of the theory, i.e. the limit of validity of the perturbative approach.

Unitarity: longitudinal gauge boson scattering cross section at high energy grows with M_H .

Electroweak Equivalence Theorem:

in the high energy limit ($s \gg M_V^2$)

$$\mathcal{A}(V_L^1 \dots V_L^n \rightarrow V_L^1 \dots V_L^m) = (i)^n (-i)^m \mathcal{A}(\omega^1 \dots \omega^n \rightarrow \omega^1 \dots \omega^m) + O\left(\frac{M_V^2}{s}\right)$$

(V_L^i =longitudinal weak gauge boson; ω^i =associated Goldstone boson).

Example: $W_L^+ W_L^- \rightarrow W_L^+ W_L^-$

$$\mathcal{A}(W_L^+ W_L^- \rightarrow W_L^+ W_L^-) \sim -\frac{1}{v^2} \left(-s - t + \frac{s^2}{s - M_H^2} + \frac{t^2}{t - M_H^2} \right)$$

$$\mathcal{A}(\omega^+ \omega^- \rightarrow \omega^+ \omega^-) = -\frac{M_H^2}{v^2} \left(\frac{s}{s - M_H^2} + \frac{t}{t - M_H^2} \right)$$

\Downarrow

$$\mathcal{A}(W_L^+ W_L^- \rightarrow W_L^+ W_L^-) = \mathcal{A}(\omega^+ \omega^- \rightarrow \omega^+ \omega^-) + O\left(\frac{M_W^2}{s}\right)$$

Using partial wave decomposition:

$$\mathcal{A} = 16\pi \sum_{l=0}^{\infty} (2l+1) P_l(\cos\theta) a_l$$

$$\frac{d\sigma}{d\Omega} = \frac{1}{64\pi^2 s} |\mathcal{A}^2| \longrightarrow \sigma = \frac{16\pi}{s} \sum_{l=0}^{\infty} (2l+1) |a_l|^2 = \frac{1}{s} \text{Im} [\mathcal{A}(\theta=0)]$$

⇓

$$\boxed{|a_l|^2 = \text{Im}(a_l)} \longrightarrow \boxed{|\text{Re}(a_l)| \leq \frac{1}{2}}$$

Most constraining condition for $W_L^+ W_L^- \rightarrow W_L^+ W_L^-$ from

$$a_0(\omega^+ \omega^- \rightarrow \omega^+ \omega^-) = -\frac{M_H^2}{16\pi v^2} \left[2 + \frac{M_H^2}{s - M_H^2} - \frac{M_H^2}{s} \log \left(1 + \frac{s}{M_H^2} \right) \right] \xrightarrow{s \gg M_H^2} -\frac{M_H^2}{8\pi v^2}$$

$$\boxed{|\text{Re}(a_0)| < \frac{1}{2}} \longrightarrow \boxed{M_H < 870 \text{ GeV}}$$

Best constraint from coupled channels ($2W_L^+ W_L^- + Z_L Z_L$):

$$a_0 \xrightarrow{s \gg M_H^2} -\frac{5M_H^2}{32\pi v^2} \longrightarrow \boxed{M_H < 780 \text{ GeV}}$$

Observe that: if there is no Higgs boson, i.e. $M_H \gg s$:

$$a_0(\omega^+\omega^- \rightarrow \omega^+\omega^-) \xrightarrow{M_H^2 \gg s} -\frac{s}{32\pi v^2}$$

Imposing the unitarity constraint \longrightarrow $\sqrt{s_c} < 1.8 \text{ TeV}$

Most restrictive constraint \longrightarrow $\sqrt{s_c} < 1.2 \text{ TeV}$



New physics expected at the TeV scale

Exciting !!

this is the range of energies of both Tevatron and LHC

Triviality: a $\lambda\phi^4$ theory cannot be perturbative at all scales unless $\lambda=0$.

In the SM the **scale evolution of λ** is more complicated:

$$32\pi^2 \frac{d\lambda}{dt} = 24\lambda^2 - (3g'^2 + 9g^2 - 24y_t^2)\lambda + \frac{3}{8}g'^4 + \frac{3}{4}g'^2g^2 + \frac{9}{8}g^4 - 24y_t^4 + \dots$$

($t = \ln(Q^2/Q_0^2)$, $y_t = m_t/v \rightarrow$ top quark Yukawa coupling).

Still, **for large λ** (\leftrightarrow large M_H) the first term dominates and (at 1-loop):

$$\lambda(Q) = \frac{\lambda(Q_0)}{1 - \frac{3}{4\pi^2}\lambda(Q_0)\ln\left(\frac{Q^2}{Q_0^2}\right)}$$

when Q grows

\longrightarrow

$\lambda(Q)$ hits a pole \rightarrow triviality

Imposing that $\lambda(Q)$ is finite, gives a scale dependent bound on M_H :

$$\frac{1}{\lambda(\Lambda)} > 0 \longrightarrow M_H^2 < \frac{8\pi^2 v^2}{3 \log\left(\frac{\Lambda^2}{v^2}\right)}$$

where we have set $Q \rightarrow \Lambda$ and $Q_0 \rightarrow v$.

Vacuum stability: $\lambda(Q) > 0$

For small λ (\leftrightarrow small M_H) the last term in $d\lambda/dt = \dots$ dominates and:

$$\lambda(\Lambda) = \lambda(v) - \frac{3}{4\pi^2} y_t^4 \log \left(\frac{\Lambda^2}{v^2} \right)$$

from where a first rough lower bound is derived:

$$\lambda(\Lambda) > 0 \quad \longrightarrow \quad M_H^2 > \frac{3v^2}{2\pi^2} y_t^4 \log \left(\frac{\Lambda^2}{v^2} \right)$$

More accurate analyses use 2-loop renormalization group improved V_{eff} .

Fine-tuning: M_H is unstable to ultraviolet corrections

$$M_H^2 = (M_H^0)^2 + \frac{g^2}{16\pi^2} \Lambda^2 \cdot \text{constant} + \text{higher orders}$$

$M_H^0 \rightarrow$ fundamental parameter of the SM

$\Lambda \rightarrow$ UV-cutoff scale

Unless $\Lambda \simeq$ EW-scale, **fine-tuning** is required to get $M_H \simeq$ EW-scale.

More generally, the all order calculation of V_{eff} would give:

$$\bar{\mu}^2 = \mu^2 + \Lambda^2 \sum_{n=0}^{\infty} c_n(\lambda_i) \log^n(\Lambda/Q)$$

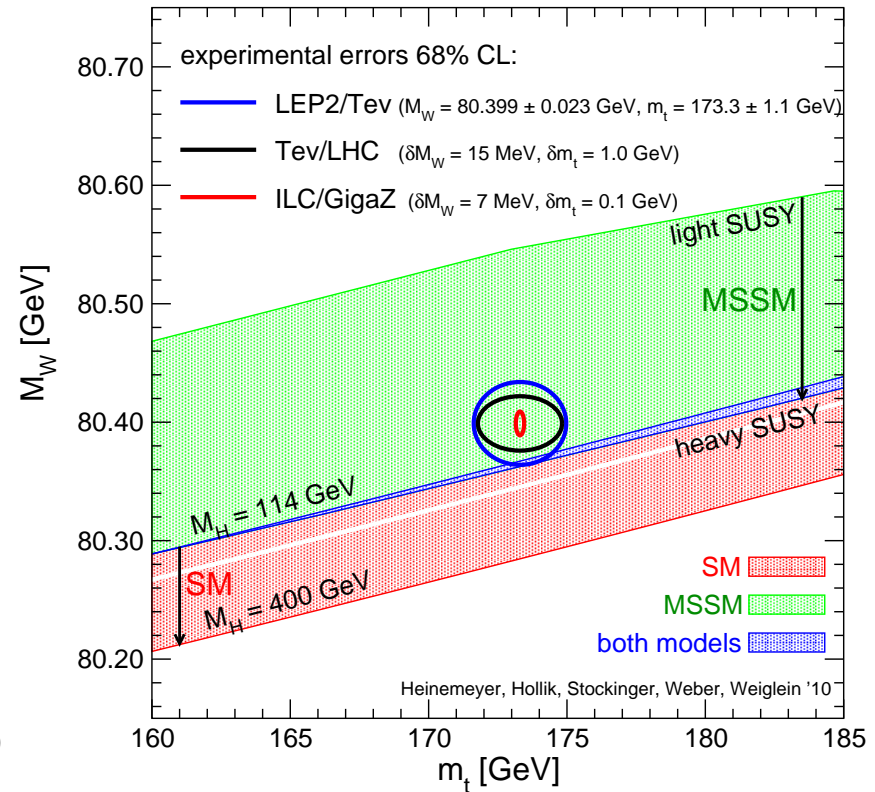
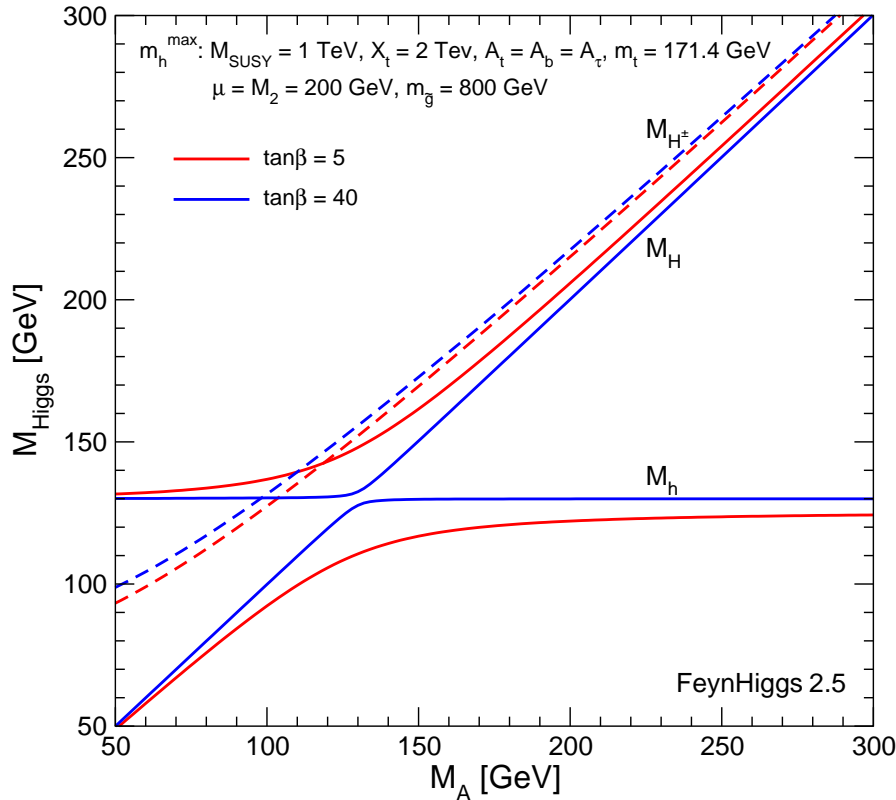
Veltman condition: the absence of large quadratic corrections is guaranteed by:

$$\sum_{n=0}^{\infty} c_n(\lambda_i) \log^n(\Lambda/M_H) = 0 \quad \text{or better} \quad \sum_{n=0}^{n_{max}} c_n(\lambda_i) \log^n(\Lambda/M_H) < \frac{v^2}{\Lambda^2}$$

where: $n_{max} = 0, 1, 2 \rightarrow \Lambda \simeq 2, 15, 50$ TeV.

Beyond SM: new physics at the TeV scale can be a better fit

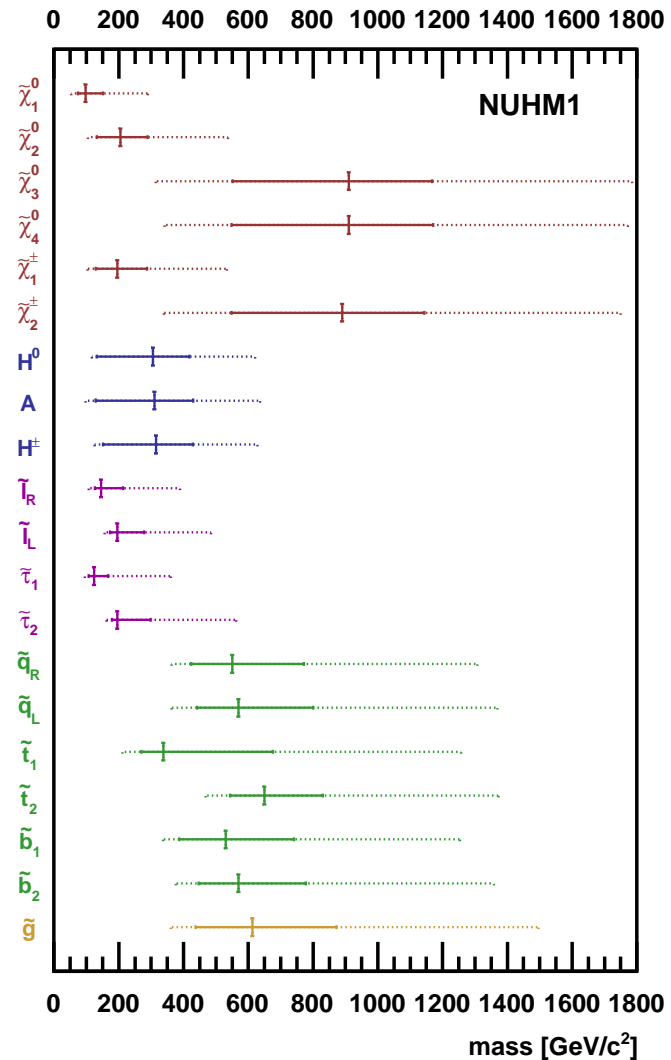
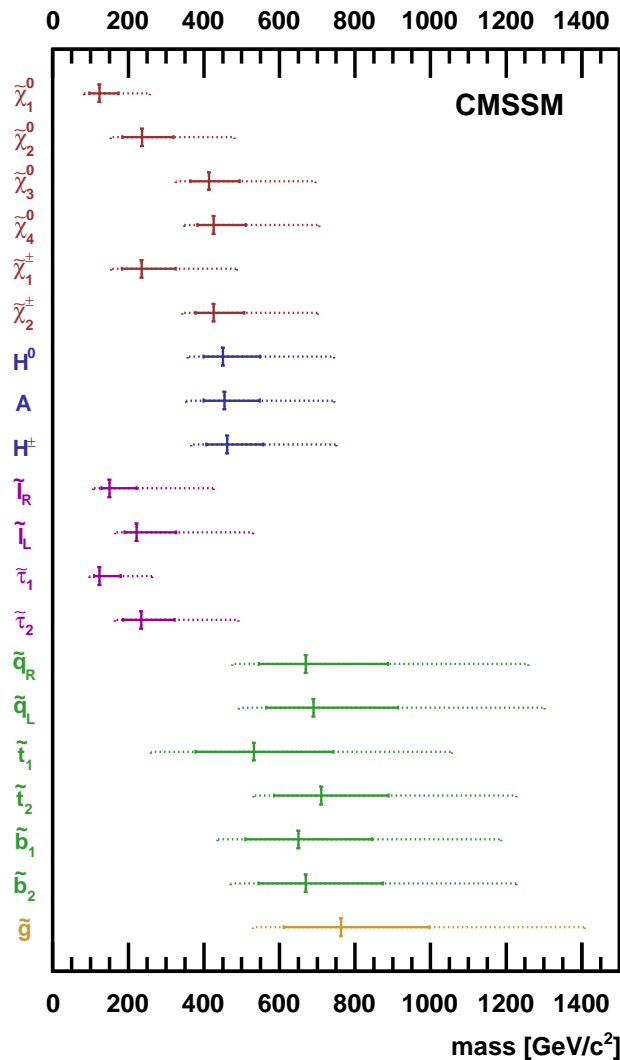
Ex. 1: MSSM



- ▶ a light scalar Higgs boson, along with a heavier scalar, a pseudoscalar and a charged scalar;
- ▶ similar although less constrained pattern in any 2HDM;
- ▶ MSSM main uncertainty: unknown masses of SUSY particles.
- ▶ precise measurement of mass spectrum and couplings will be crucial.

... mass spectrum at a glance ...

(MasterCode by Buchmüller et al., '09)



- ▶ CMSSM/NUHM1 (different choice of soft SUSY breaking mass terms);
- ▶ all available data (exp.) and all known corrections (th.) included in fit;
- ▶ most masses accessible to early LHC, several within reach of ILC.

The Higgs bosons of the MSSM: example of 2HDM

Two complex $SU(2)_L$ doublets, with hypercharge $Y = \pm 1$:

$$\Phi_u = \begin{pmatrix} \phi_u^+ \\ \phi_u^0 \end{pmatrix}, \quad \Phi_d = \begin{pmatrix} \phi_d^0 \\ \phi_d^- \end{pmatrix}$$

and (super)potential (Higgs part only):

$$\begin{aligned} V_H &= (|\mu|^2 + m_u^2)|\Phi_u|^2 + (|\mu|^2 + m_d^2)|\Phi_d|^2 - \mu B \epsilon_{ij} (\Phi_u^i \Phi_d^j + h.c.) \\ &+ \frac{g^2 + g'^2}{8} (|\Phi_u|^2 - |\Phi_d|^2)^2 + \frac{g^2}{2} |\Phi_u^\dagger \Phi_d|^2 \end{aligned}$$

The EW symmetry is spontaneously broken by choosing:

$$\langle \Phi_u \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_u \end{pmatrix}, \quad \langle \Phi_d \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} v_d \\ 0 \end{pmatrix}$$

normalized to preserve the SM relation:

$$M_W^2 = g^2(v_u^2 + v_d^2)/4 = g^2 v^2/4.$$

Five physical scalar/pseudoscalar degrees of freedom:

$$h^0 = -(\sqrt{2}\text{Re}\Phi_d^0 - v_d) \sin \alpha + (\sqrt{2}\text{Re}\Phi_u^0 - v_u) \cos \alpha$$

$$H^0 = (\sqrt{2}\text{Re}\Phi_d^0 - v_d) \cos \alpha + (\sqrt{2}\text{Re}\Phi_u^0 - v_u) \sin \alpha$$

$$A^0 = \sqrt{2} (\text{Im}\Phi_d^0 \sin \beta + \text{Im}\Phi_u^0 \cos \beta)$$

$$H^\pm = \Phi_d^\pm \sin \beta + \Phi_u^\pm \cos \beta$$

where $\boxed{\tan \beta = v_u/v_d}$.

All masses can be expressed (at tree level) in terms of $\boxed{\tan \beta}$ and M_A :

$$M_{H^\pm}^2 = M_A^2 + M_W^2$$

$$M_{H,h}^2 = \frac{1}{2} \left(M_A^2 + M_Z^2 \pm ((M_A^2 + M_Z^2)^2 - 4M_Z^2 M_A^2 \cos^2 2\beta)^{1/2} \right)$$

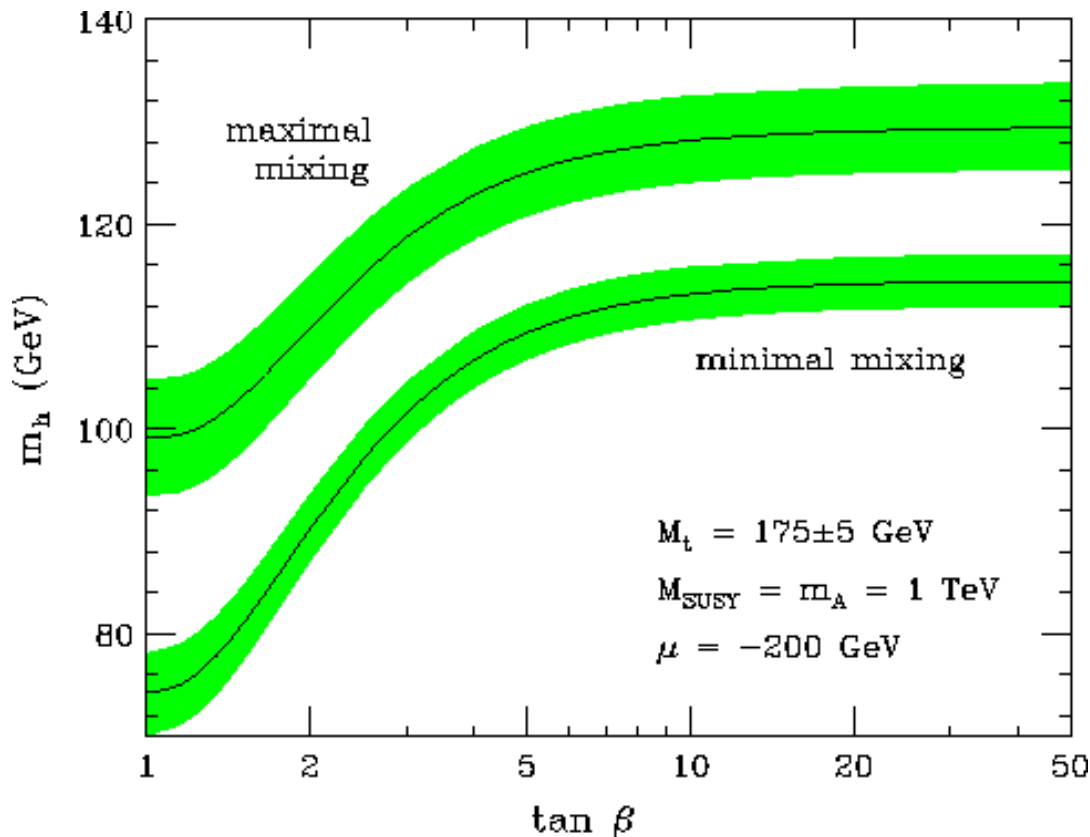
Notice: tree level upper bound on M_h : $\boxed{M_h^2 \leq M_Z^2 \cos 2\beta \leq M_Z^2}$!

Higgs masses greatly modified by radiative corrections.

In particular, the upper bound on M_h becomes:

$$M_h^2 \leq M_Z^2 + \frac{3g^2 m_t^2}{8\pi^2 M_W^2} \left[\log \left(\frac{M_S^2}{m_t^2} \right) + \frac{X_t^2}{M_S^2} \left(1 - \frac{X_t^2}{12M_S^2} \right) \right]$$

where $M_S \equiv (M_{\tilde{t}_1}^2 + M_{\tilde{t}_2}^2)/2$ while X_t is the top squark mixing parameter:



$$\begin{pmatrix} M_{Q_t}^2 + m_t^2 + D_L^t & m_t X_t \\ m_t X_t & M_{R_t}^2 + m_t^2 + D_R^t \end{pmatrix}$$

with $X_t \equiv A_t - \mu \cot \beta$.

$$D_L^t = (1/2 - 2/3 \sin \theta_W) M_Z^2 \cos 2\beta$$

$$D_R^t = 2/3 \sin^2 \theta_W M_Z^2 \cos 2\beta$$

Higgs boson couplings to SM gauge bosons:

Some phenomenologically important ones:

$$g_{hVV} = g_V M_V \sin(\beta - \alpha) g^{\mu\nu} \quad , \quad g_{HVV} = g_V M_V \cos(\beta - \alpha) g^{\mu\nu}$$

where $g_V = 2M_V/v$ for $V = W, Z$, and

$$g_{hAZ} = \frac{g \cos(\beta - \alpha)}{2 \cos \theta_W} (p_h - p_A)^\mu \quad , \quad g_{HAZ} = -\frac{g \sin(\beta - \alpha)}{2 \cos \theta_W} (p_H - p_A)^\mu$$

Notice: $g_{AZZ} = g_{AWW} = 0$, $g_{H^\pm ZZ} = g_{H^\pm WW} = 0$

Decoupling limit: $M_A \gg M_Z \longrightarrow \begin{cases} M_h \simeq M_h^{max} \\ M_H \simeq M_{H^\pm} \simeq M_A \end{cases}$

$$\cos^2(\beta - \alpha) \simeq \frac{M_Z^4 \sin^2 4\beta}{M_A^4} \longrightarrow \begin{cases} \cos(\beta - \alpha) \rightarrow 0 \\ \sin(\beta - \alpha) \rightarrow 1 \end{cases}$$

The only low energy Higgs is $h \simeq H_{SM}$.

Higgs boson couplings to quarks and leptons:

Yukawa type couplings, Φ_u to up-component and Φ_d to down-component of $SU(2)_L$ fermion doublets. Ex. (3rd generation quarks):

$$\mathcal{L}_{Yukawa} = h_t [\bar{t}P_L t \Phi_u^0 - \bar{t}P_L b \Phi_u^+] + h_b [\bar{b}P_L b \Phi_d^0 - \bar{b}P_L t \Phi_d^-] + \text{h.c.}$$

and similarly for leptons. The corresponding couplings can be expressed as ($y_t, y_b \rightarrow \text{SM}$):

$$\begin{aligned} g_{ht\bar{t}} &= \frac{\cos \alpha}{\sin \beta} y_t = [\sin(\beta - \alpha) + \cot \beta \cos(\beta - \alpha)] y_t \\ g_{hb\bar{b}} &= -\frac{\sin \alpha}{\cos \beta} y_b = [\sin(\beta - \alpha) - \tan \beta \cos(\beta - \alpha)] y_b \\ g_{Ht\bar{t}} &= \frac{\sin \alpha}{\sin \beta} y_t = [\cos(\beta - \alpha) - \cot \beta \sin(\beta - \alpha)] y_t \\ g_{Hb\bar{b}} &= \frac{\cos \alpha}{\cos \beta} y_b = [\cos(\beta - \alpha) + \tan \beta \sin(\beta - \alpha)] y_b \\ g_{At\bar{t}} &= \cot \beta y_t \quad , \quad g_{Ab\bar{b}} = \tan \beta y_b \\ g_{H^\pm t\bar{b}} &= \frac{g}{2\sqrt{2}M_W} [m_t \cot \beta (1 + \gamma_5) + m_b \tan \beta (1 - \gamma_5)] \end{aligned}$$

Notice: consistent decoupling limit behavior.

Higgs couplings modified by radiative corrections

Most important effects:

- **Corrections to $\cos(\beta - \alpha)$** : crucial in decoupling behavior.

$$\cos(\beta - \alpha) = K \left[\frac{M_Z^2 \sin 4\beta}{2M_A^2} + \mathcal{O} \left(\frac{M_Z^4}{M_A^4} \right) \right]$$

where

$$K \equiv 1 + \frac{\delta\mathcal{M}_{11}^2 - \delta\mathcal{M}_{22}^2}{2M_Z^2 \cos 2\beta} - \frac{\delta\mathcal{M}_{12}^2}{M_Z^2 \sin 2\beta}$$

$\delta\mathcal{M}_{ij}$ \longrightarrow corrections to the CP-even scalar mass matrix.

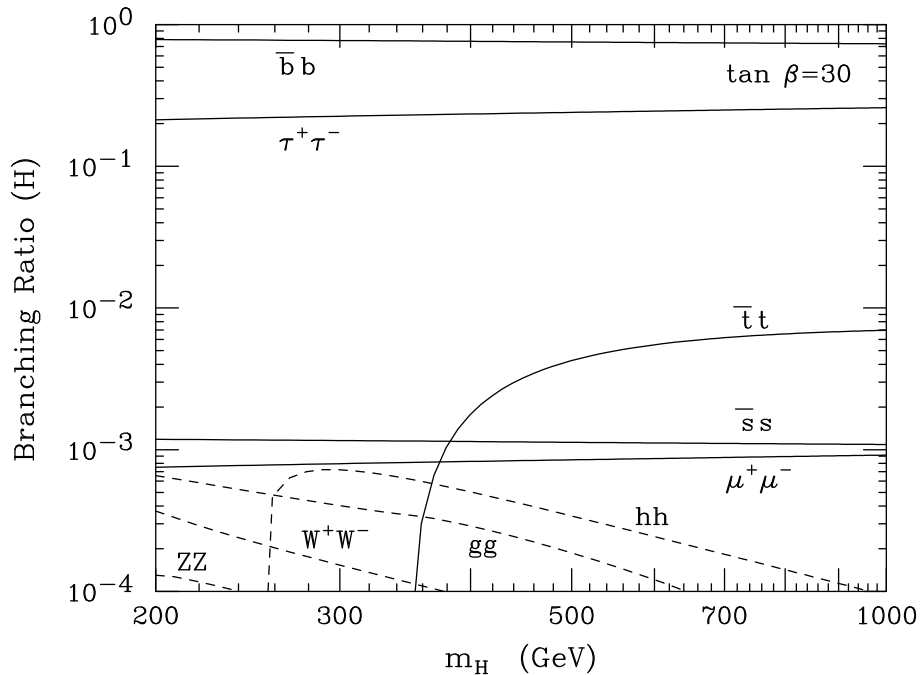
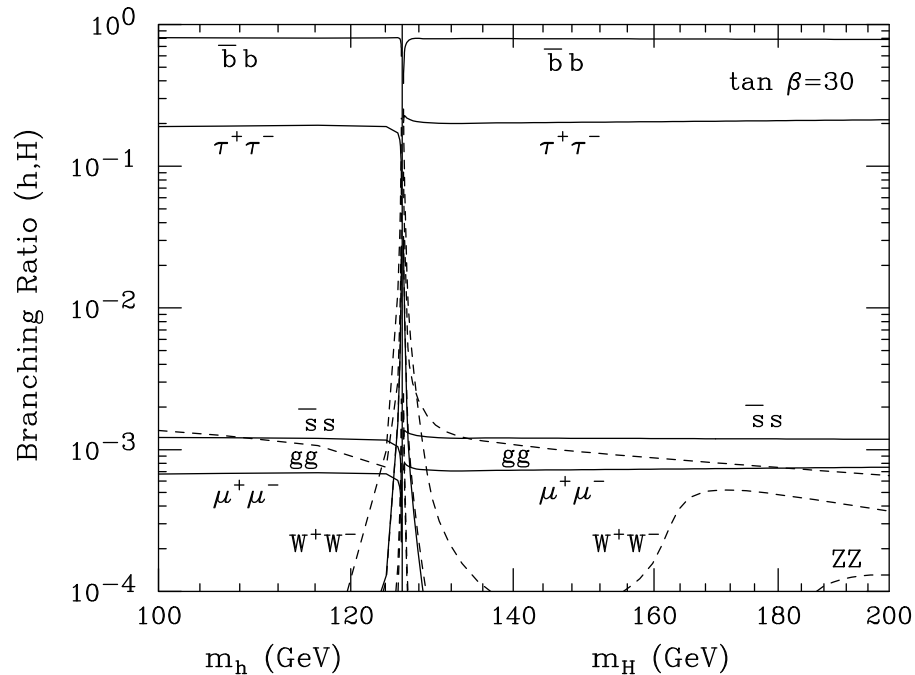
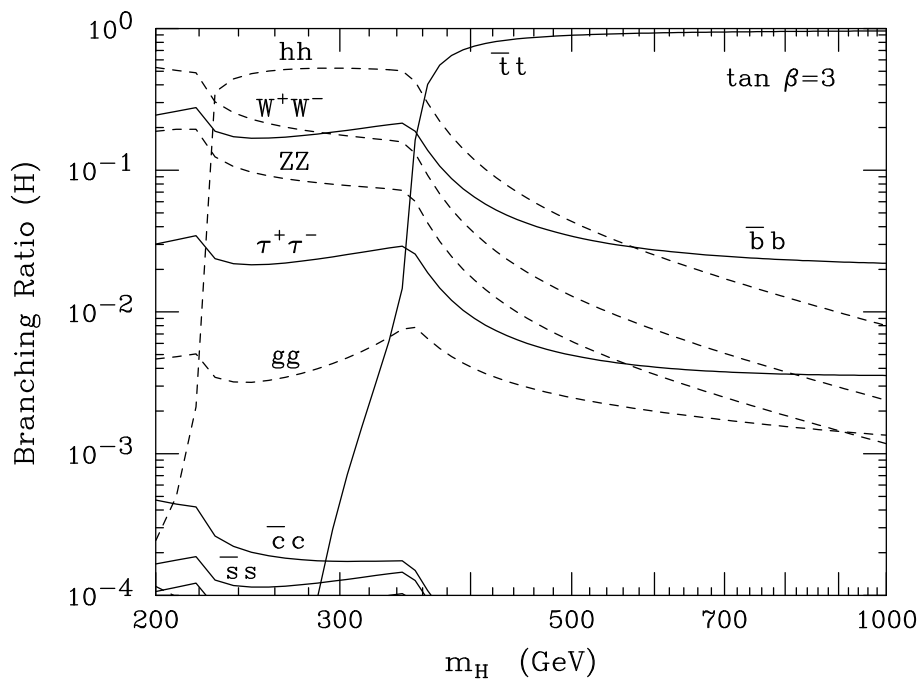
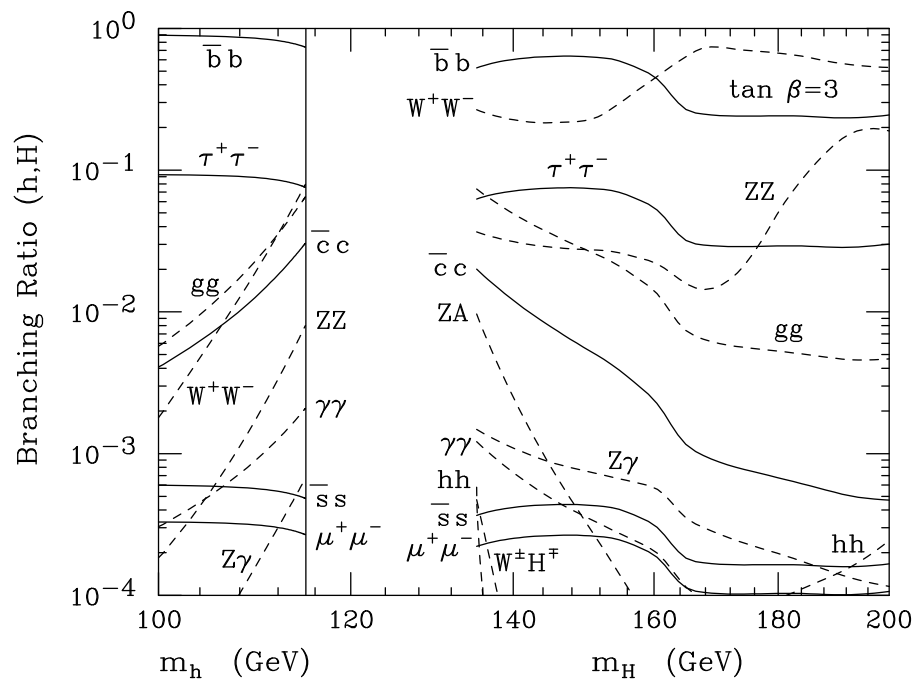
- **Corrections to 3rd generation Higgs-fermion Yukawa couplings.**

$$-\mathcal{L}_{eff} = \epsilon_{ij} \left[(h_b + \delta h_b) \bar{b}_R H_d^i Q_L^j + (h_t + \delta h_t) \bar{t}_R H_u^j Q_L^i \right] + \Delta h_t \bar{t}_R Q_L^k H_d^{k*} + \Delta h_b \bar{b}_R Q_L^k H_u^{k*} + \text{h.c.}$$

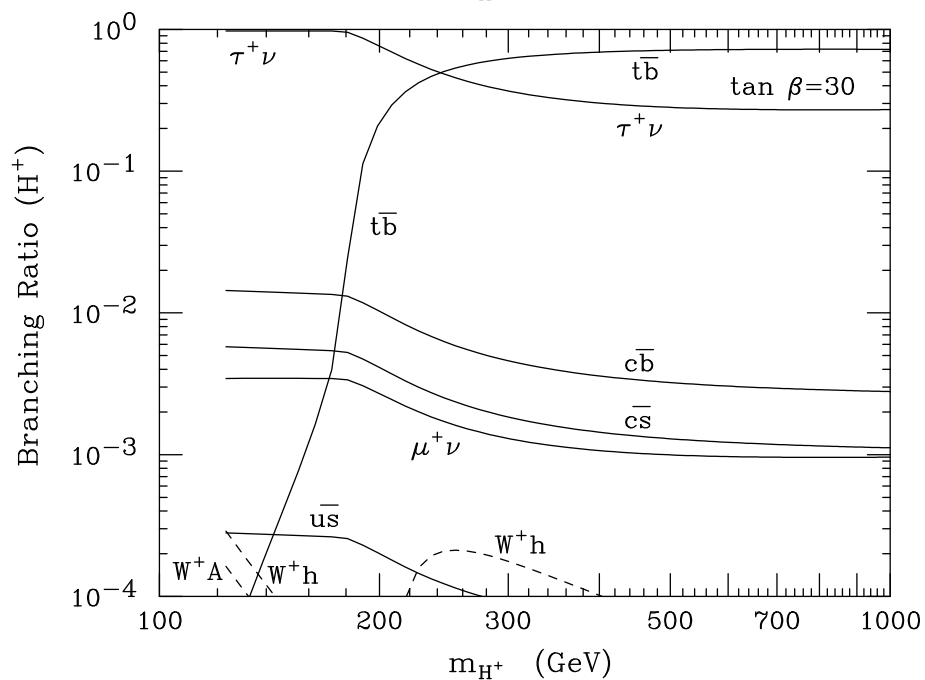
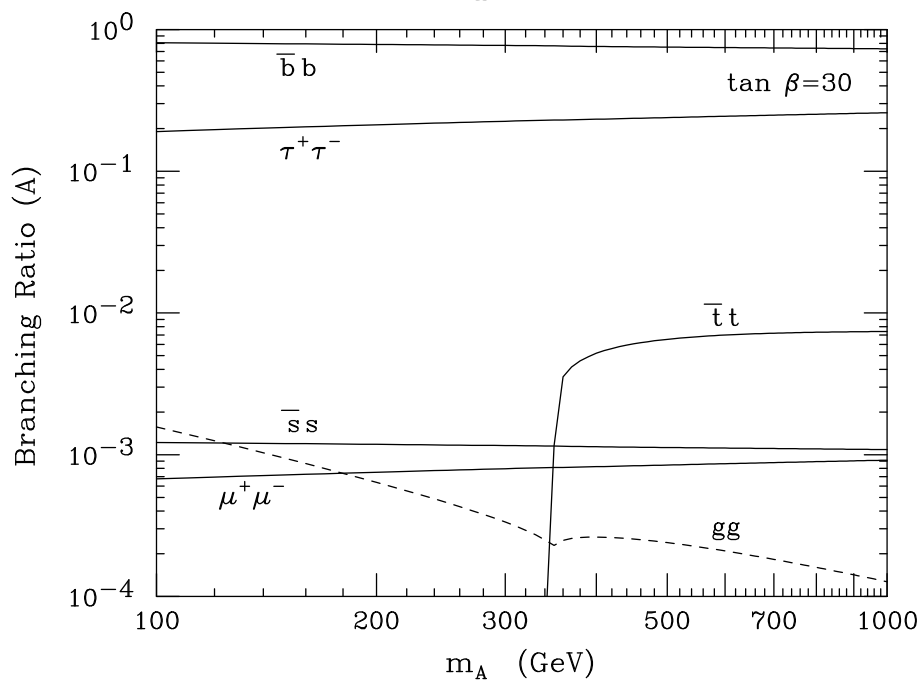
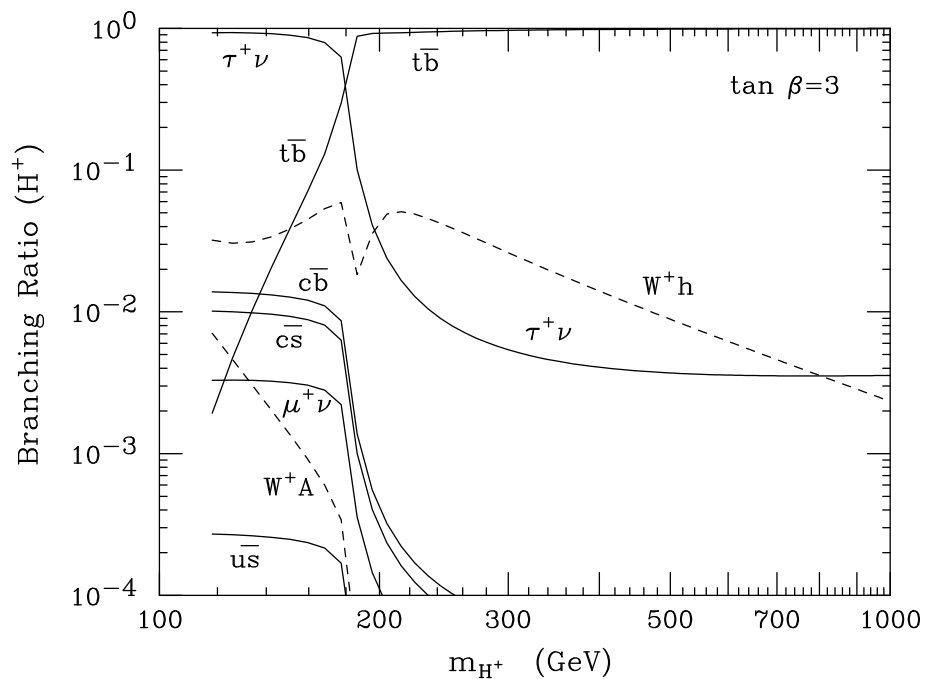
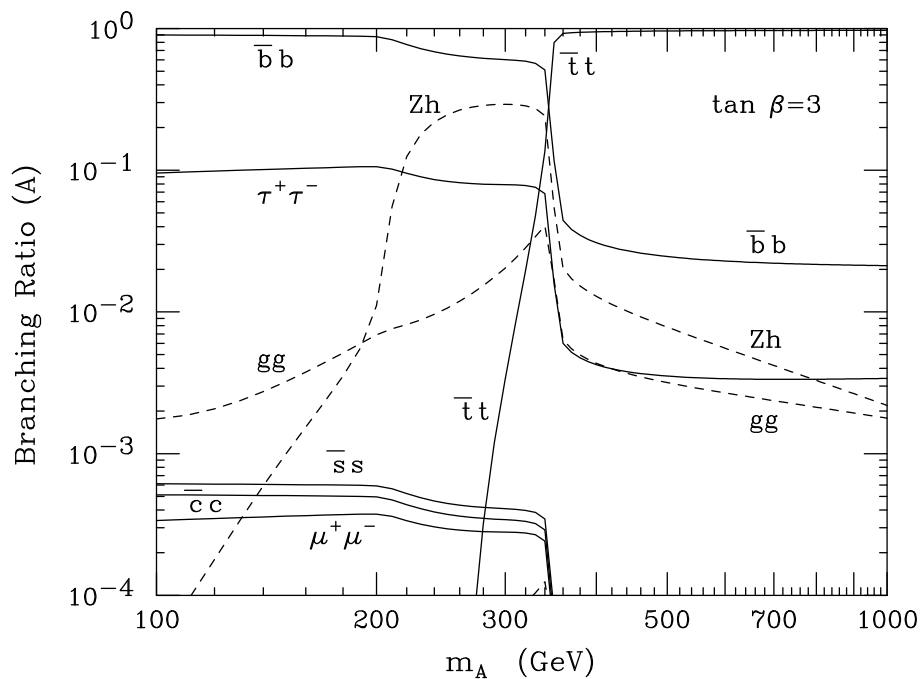
$$m_b = \frac{h_b v}{\sqrt{2}} \cos \beta \left(1 + \frac{\delta h_b}{h_b} + \frac{\Delta h_b \tan \beta}{h_b} \right) \equiv \frac{h_b v}{\sqrt{2}} \cos \beta (1 + \Delta_b)$$

$$m_t = \frac{h_t v}{\sqrt{2}} \sin \beta \left(1 + \frac{\delta h_t}{h_t} + \frac{\Delta h_t \tan \beta}{h_t} \right) \equiv \frac{h_t v}{\sqrt{2}} \sin \beta (1 + \Delta_t)$$

MSSM Higgs boson branching ratios, possible scenarios:

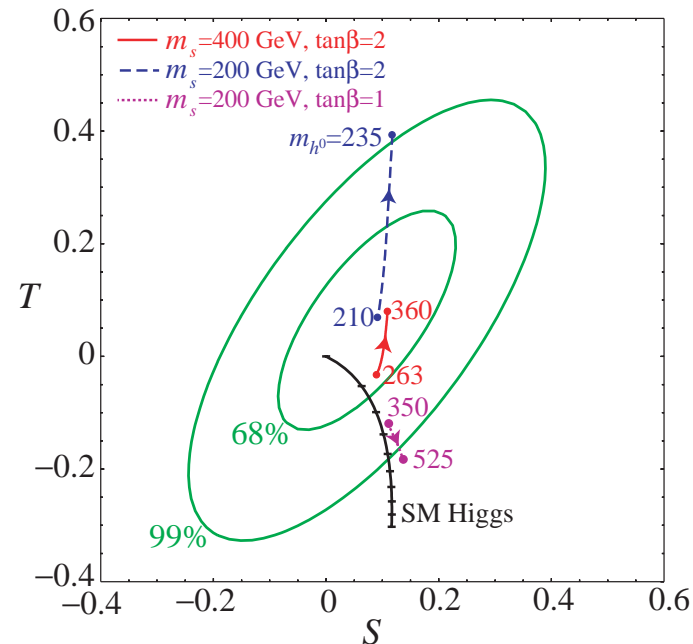
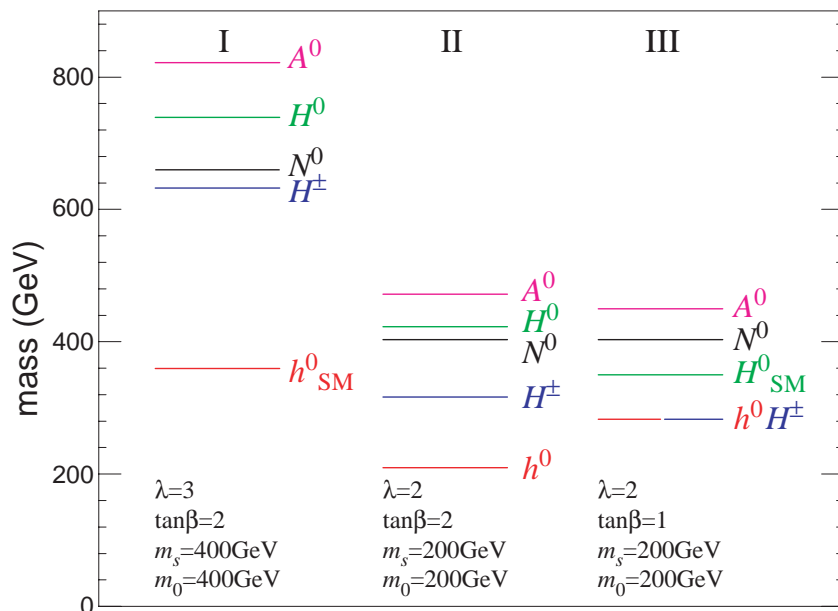


MSSM Higgs boson branching ratios, possible scenarios:



Beyond SM: new physics at the TeV scale can be a better fit

Ex. 2: “Fat Higgs” models



(Harnik, Kribs, Larson, and Murayama, PRD 70 (2004) 015002)

- ▷ supersymmetric theory of a composite Higgs boson;
- ▷ moderately heavy lighter scalar Higgs boson, along with a heavier scalar, a pseudoscalar and a charged scalar;
- ▷ consistent with EW precision measurements without fine tuning.