

## Special and General Relativity (PHZ 4601/5606) Fall 2017 Solutions

### Set 4

#### 12. Proper time under rotation.

$$c^2 d\tau^2 = c^2 dt^2 - \omega^2 R^2 dt^2 \Rightarrow \frac{d\tau}{dt} = \sqrt{1 - \frac{\omega^2 R^2}{c^2}}.$$

Now  $g = v^2/R = \omega^2 R$ , where  $g = 9.81 [m/s^2]$  and  $v$  is the velocity of the satellite. For  $R = (6,378 + 160) [km]$  (equatorial radius plus 160 [km]) we get  $\omega = 0.0012249 [s^{-1}]$ , which corresponds to a period of  $T = 2\pi/\omega = 85.5 [minutes]$ . For the time dilation we find

$$1 - \frac{d\tau}{dt} = \frac{1}{2} \left( \frac{\omega R}{c} \right)^2 = 0.356321 \times 10^{-9}.$$

Compared to this the time dilation of a clock on the equator is with

$$1 - \frac{d\tau}{dt} = \frac{1}{2} \left( \frac{\omega R}{c} \right)^2 = 0.119517 \times 10^{-11}$$

negligible. For gravity we obtain along the lines of the previous problem 4

$$1 - \frac{d\tau}{dt} = -0.17422 \times 10^{-10},$$

which reduces the the SR effect by about 5%.