

ADVANCED DYNAMICS — PHY 4241/5227

HOME AND CLASS WORK – SET 7

(February 12, 2010)

(23) Assume a central potential and $\vec{L} \neq 0$.

1. Use initial conditions in the CM system, \vec{r}_0 and \vec{v}_0 to define unit vectors \hat{e}_1 , \hat{e}_2 and \hat{e}_3 of a Cartesian coordinates system in which

$$L = |\vec{L}| = \mu r^2 \dot{\phi}, \quad \phi_0 = 0, \quad \dot{\phi}_0 > 0$$

holds (ϕ azimuth angle). Due February 12 in class (4 points).

2. Express the total energy E_{cm} in terms of r and ϕ and their derivatives (i.e., get rid of \vec{v}). Substitute the conserved angular momentum L into your expression. Combine all terms, which contain r but not \dot{r} , into the definition of an effective potential $U_{\text{eff}}(r)$. Due February 12 in class (4 points).

(24) Use (in arbitrary units) the parameters $\alpha = \mu = L = 1$ to plot the potential and the effective potential of the Kepler problem

$$U = -\frac{\alpha}{r} \quad \text{and} \quad U_{\text{eff}}(r) = U(r) + \frac{L^2}{2\mu r^2}$$

in the range $0 < r \leq 10$ with the ordinate restricted to the interval $[-1, 0.6]$. Describe the motion for $E_{\text{cm}} = U_{\text{eff}}^{\text{min}}$, where $U_{\text{eff}}^{\text{min}} = U_{\text{eff}}(r^{\text{min}})$ is the minimum of the effective potential. Calculate r^{min} as function of L , μ and α and give then the numerical value. Due February 15 before class (8 points).

(25a) Separate variables in the equation

$$E_{\text{cm}} = \frac{\mu}{2} \dot{r}^2 + U_{\text{eff}}(r)$$

to derive an integral equation for $t(r)$. Due February 15 in class (4 points).

(25b) Use angular momentum conservation and the separation of variables in the previous problem to derive an integral equation for $\phi(r)$. Due February 15 in class (4 points).