

## Electrodynamics A (PHY 5346) Fall 2016 Solutions

### Set 4:

(12) Exercise: End of Spacetrip.

We use energy-momentum conservation

$$0 = dp = \begin{pmatrix} dp_r^0 \\ dp_r^1 \\ 0 \\ 0 \end{pmatrix} + \begin{pmatrix} dp_e^0 \\ dp_e^1 \\ 0 \\ 0 \end{pmatrix},$$

where the subscript  $r$  stands for rocket and  $e$  for exhaust. In the temporary rest frame of the rocket we have

$$dp_e^1 = -v dm, \quad dp_r^1 = m du = m g d\tau,$$

where  $v$  is the velocity of the exhaust,  $dm$  the infinitesimal change of rest mass of the rocket and  $du$  the infinitesimal velocity of the rocket. Note that the exhaust may have no rest mass and our equation for  $dp_e^1$  (no  $\gamma$ ) is not obvious. It follows from

$$dp_e^0 = -dp_r^0 = c dm \quad \text{and} \quad -\beta = \frac{dp_e^1}{dp_e^0}.$$

With this definition  $\beta = v/c$  is positive, because  $dp_e^1$  is negative when we choose  $dp_r^1$  positive. Now, separation of variables gives

$$\frac{dm}{m} = -\frac{g}{v} d\tau \Rightarrow \int_{m_0}^{m(\tau)} \frac{dm}{m} = -\frac{g}{v} \int_0^\tau d\tau' \Rightarrow \ln\left(\frac{m(\tau)}{m_0}\right) = -\frac{g\tau}{v}.$$

(1) The mass of the spaceship decreases with its proper time according to

$$m(\tau) = m_0 \exp\left(-\frac{g\tau}{v}\right).$$

(2) With  $v = 0.66c$  and after a trip of twenty years the remaining fraction of the mass is  $m(20 \text{ years})/m_0 = 2.68 \times 10^{-14}$ .

(3) With  $v = c$  it becomes  $m(20 \text{ years})/m_0 = 1.1 \times 10^{-9}$ .

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(13) Exercise: Redshift.

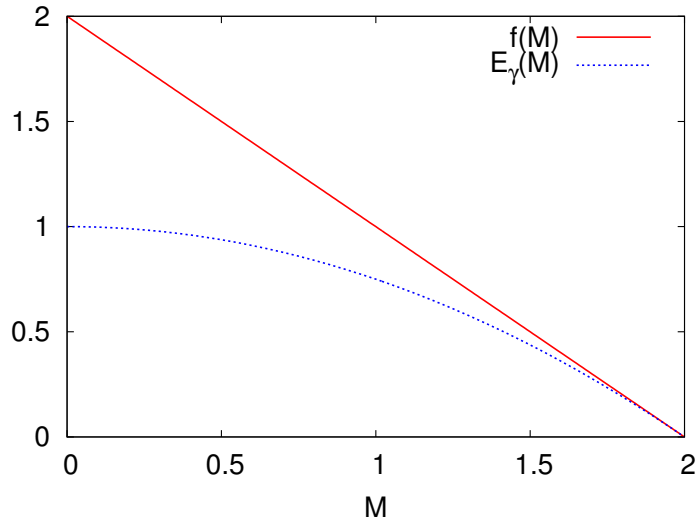
The equation for the redshift is

$$\begin{aligned}\lambda' &= \lambda \sqrt{\frac{1+\beta}{1-\beta}} \Rightarrow \left(\frac{\lambda'}{\lambda}\right)^2 = \frac{1+\beta}{1-\beta} \\ \left(\frac{\lambda'}{\lambda}\right)^2 - \beta \left(\frac{\lambda'}{\lambda}\right)^2 &= 1 + \beta \\ \left(\frac{\lambda'}{\lambda}\right)^2 - 1 &= \beta \left[1 + \left(\frac{\lambda'}{\lambda}\right)^2\right] \Rightarrow \beta = \frac{(\lambda'/\lambda)^2 - 1}{(\lambda'/\lambda)^2 + 1}.\end{aligned}$$

With  $\lambda' = (729.2 \text{ [nm]}) m^2 / (m^2 - 4)$  and  $\lambda = (364.56 \text{ [nm]}) m^2 / (m^2 - 4)$  we find

$$\beta = \frac{v}{c} = 0.6.$$

(14) Exercise: Relativistic energy-momentum conservation.

Fig. 0.1 Comparison of  $f(M) = 2m - M$  with photon energy  $E_\gamma(M)$ .

1. The 4-vector energy-momentum conservation reads

$$\begin{pmatrix} m \\ 0 \\ 0 \\ 0 \end{pmatrix} + \begin{pmatrix} m \\ 0 \\ 0 \\ 0 \end{pmatrix} = \begin{pmatrix} \sqrt{M^2 + p^2} \\ -p \\ 0 \\ 0 \end{pmatrix} + \begin{pmatrix} E_\gamma \\ p_\gamma \\ 0 \\ 0 \end{pmatrix}$$

Therefore (eliminating  $p$  via  $p = p_\gamma = E_\gamma$ ),

$$2m - E_\gamma = \sqrt{M^2 + E_\gamma^2},$$

$$(2m - E_\gamma)^2 = 4m^2 - 4m E_\gamma + E_\gamma^2 = M^2 + E_\gamma^2,$$

$$E_\gamma = \frac{4m^2 - M^2}{4m}.$$

2. The requested  $E_\gamma(M)$  values are:  $E_\gamma(0) = m$ ,

$$E_\gamma\left(\frac{m}{2}\right) = \frac{15m}{16}, \quad E_\gamma(m) = \frac{3m}{4}, \quad E_\gamma(\sqrt{2}m) = \frac{m}{2}, \quad E_\gamma(\sqrt{3}m) = \frac{m}{4}.$$

3. The sketch is given in the figure.

(15) **Electromagnetic field tensor in  $\vec{E}$  and  $\vec{B}$  fields.**

A. From the  $J^0$  component we get

$$\partial_0 F^{00} + \partial_1 F^{10} + \partial_2 F^{20} + \partial_3 F^{30} = \frac{4\pi}{c} J^0 = 4\pi \rho = \nabla \cdot \vec{E} = \partial_1 E^1 + \partial_2 E^2 + \partial_3 E^3.$$

Therefore,  $F^{00} = 0$  by anti-symmetry,  $F^{10} = E^1$ ,  $F^{20} = E^2$  and  $F^{30} = E^3$ .

From the  $J^1$  component we get

$$\partial_0 F^{01} + \partial_1 F^{11} + \partial_2 F^{21} + \partial_3 F^{31} = \frac{4\pi}{c} J^1 = \partial_2 B^3 - \partial_3 B^2 - \partial_0 E^1.$$

Therefore,  $F^{01} = -E^1$  consistent with  $F^{10} = E^1$ ,  $F^{11} = 0$  by anti-symmetry,  $F^{21} = B^3$  and  $F^{31} = -B^2$ .

From the  $J^2$  component we get

$$\partial_0 F^{02} + \partial_1 F^{12} + \partial_2 F^{22} + \partial_3 F^{32} = \frac{4\pi}{c} J^2 = \partial_3 B^1 - \partial_1 B^3 - \partial_0 E^2.$$

Therefore,  $F^{02} = -E^2$ ,  $F^{12} = -B^3$  both consistent with anti-symmetry,  $F^{22} = 0$  by anti-symmetry and (new)  $F^{32} = B^1$ .

Using anti-symmetry all components are now determined and the  $F^{\alpha\beta}$  field tensor reads

$$(F^{\alpha\beta}) = \begin{pmatrix} 0 & -E^1 & -E^2 & -E^3 \\ E^1 & 0 & -B^3 & B^2 \\ E^2 & B^3 & 0 & -B^1 \\ E^3 & -B^2 & B^1 & 0 \end{pmatrix}.$$

The remaining equations from  $\beta = 3$  can be used for consistency checks.

B. For  $\beta = 0$ :

$$\partial_\alpha {}^*F^{\alpha 0} = \partial_1 B^1 + \partial_2 B^2 + \partial_3 B^3 = \nabla \cdot \vec{B} = 0.$$

For  $\beta = 1$ :

$$\partial_\alpha {}^*F^{\alpha 1} = -\partial_0 B^1 - \partial_2 E^3 + \partial_3 E^2 = -\partial_0 B^1 - (\nabla \times \vec{E})^1 = 0.$$

For  $\beta = 2$ :

$$\partial_\alpha {}^*F^{\alpha 2} = -\partial_0 B^2 + \partial_1 E^3 - \partial_3 E^1 = -\partial_0 B^2 - (\nabla \times \vec{E})^2 = 0.$$

For  $\beta = 3$ :

$$\partial_\alpha {}^*F^{\alpha 3} = -\partial_0 B^3 - \partial_1 E^2 + \partial_2 E^1 = -\partial_0 B^3 - (\nabla \times \vec{E})^3 = 0.$$

So,

$$\nabla \cdot \vec{B} = 0 \quad \text{and} \quad -\partial_0 \vec{B} - \nabla \times \vec{E} = 0.$$