



# Clockwork in the Heavens (Part I)

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Jan. 12, 2004

1. Large Numbers and Distances
2. Practice Quiz
3. Motion
4. Summary



# Announcements

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- Prof. Prosper will miss the next 2 weeks due to family emergency
- Course website:
  - <http://www.physics.fsu.edu/users/ProsperH/AST1002>
- Prof. Adams will continue to fill in
- Lecture notes available at:
  - <http://www.hep.fsu.edu/~tadams/courses/spr04/ast1002-2/>



# Review

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- Astronomy is the study of objects outside of the Earth's atmosphere
- This is a science class
- The scientific method always tests and retests hypotheses and develops new theories if old ones fail
- We are going to study lots of interesting stuff this semester

# Big and Small Numbers

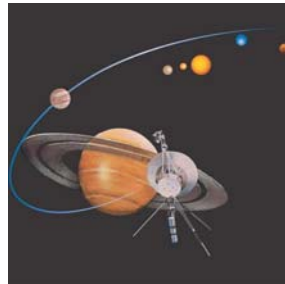
- Really big and really small numbers are hard to understand...
  - Examples, how many are 1, 10, 100, 1000, 10000, 100000, 1000000, ... ?
  - We will often use powers of 10 for large and small numbers
    - $1,000,000 = 10 \times 10 \times 10 \times 10 \times 10 \times 10 = 10^6$
  - Each factor of 10 is one order of magnitude



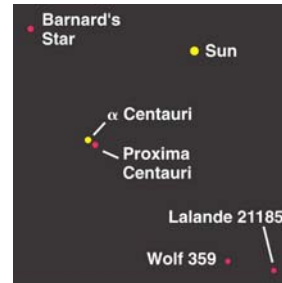
10 cm = 100 mm



10,000 km =  $10^7$  m



$10^{13}$  m = 9 lt hr



$10^{17}$  m = 10 lt yr



$10^{19}$  m =  $10^3$  lt yr



$10^{23}$  m =  $10^7$  lt yr



# Big and Small Numbers

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- Really big and really small numbers are hard to understand and work with...
  - It takes a LOT of zeros to write some numbers
    - A million billion is 1,000,000,000,000,000
  - Easier to use powers of 10
    - A million billion has 15 zeros which is  $10^{15}$
    - 5 million billion is  $5 \times 10^{15}$
  - Small numbers use negative powers of 10
    - A million billionth is  $10^{-15}$



# Measuring Distance

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## ■ Metric units

### ■ Meters (m) and kilometers (km)

- 1 meter = 3.281 feet
- Kilometer = 1000 meters =  $10^3$  meters = 0.6 miles\

## ■ Astronomical Unit (AU)

- Distance from the Sun to Earth
- 150 million kilometer

## ■ Lightyear (LY)

- Distance light travels in 1 year
- Light travels at  $3 \times 10^8$  m/s in a vacuum
  - 186,000 miles/second
  - 7 times around the Earth in a second
- So in a year, how far does light travel?



# Lightyear

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- There are about  $3 \times 10^7$  seconds in a year
  - $60 \text{ s/min} \times 60 \text{ min/hr} \times 24 \text{ hr/day} \times 365 \text{ days/year} \sim 3.15 \times 10^7 \text{ s/yr}$
- Light travels at  $3 \times 10^8 \text{ m/s}$  (or 186,000 miles/s)
- So,

$$3 \times 10^8 \text{ m/s} \times 3 \times 10^7 \text{ s} = 9 \times 10^{15} \text{ m} = 9 \times 10^{12} \text{ km}$$

$$186,000 \text{ miles/s} \times 3 \times 10^7 \text{ s} = 5.6 \times 10^{12} \text{ miles}$$



# Lightyear (cont)

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- **Actually, 1 LY =  $9.46 \times 10^{15}$  m**
  - Ten trillion kilometers
- **Sounds far, but we'll see some really far distances**
  - The nearest star is 4.3 LY away
- **It takes light one year to travel  $9.46 \times 10^{12}$  km**
  - If you were 1 LY away and flashed a light, we wouldn't see it for a year
  - Starlight we see was emitted by the stars many years ago
  - Looking into the sky is looking at what happened in the past. The further away you look, the farther back in time
- **The sun is 8 lightminutes away from us**
  - If the sun exploded right now, we wouldn't know for 8 minutes



*Wait a moment...*



# LOTS of Motion

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- Earth Spins Around Its Axis
  - Once per ???
- Earth and Moon Revolve Around Each Other
  - Once per ???
- Earth Revolves Around the Sun
  - Once per ???
- Solar System is Revolving Around the Center of the Milky Way
- The Milky Way is Moving Through Space
- Whew, do you feel dizzy?

# Earth Spins



- The Earth spins around its axis once per day (24 hr)
- When viewed from above the North Pole, the Earth rotates counterclockwise.
- Spin causes the rising and setting of the Sun (and the Moon and the stars)
- Effects many of our weather patterns including hurricanes



# North Celestial Pole

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- The Earth revolves around an axis that runs from the north celestial pole to the south celestial pole
  - Currently, the north celestial pole points towards Polaris, otherwise known as the North Star
  - There is no corresponding "South Star"
- Zenith
  - straight overhead



# The Horizon

- We can only see half the sky at any given moment
  - The other half is blocked by the Earth
- If on the North or South Pole, we always see the same half of the sky
  - It does rotate around itself
- If on the equator, we see the whole sky once per day
- In between, we see part of the sky all day long and part only some of the day



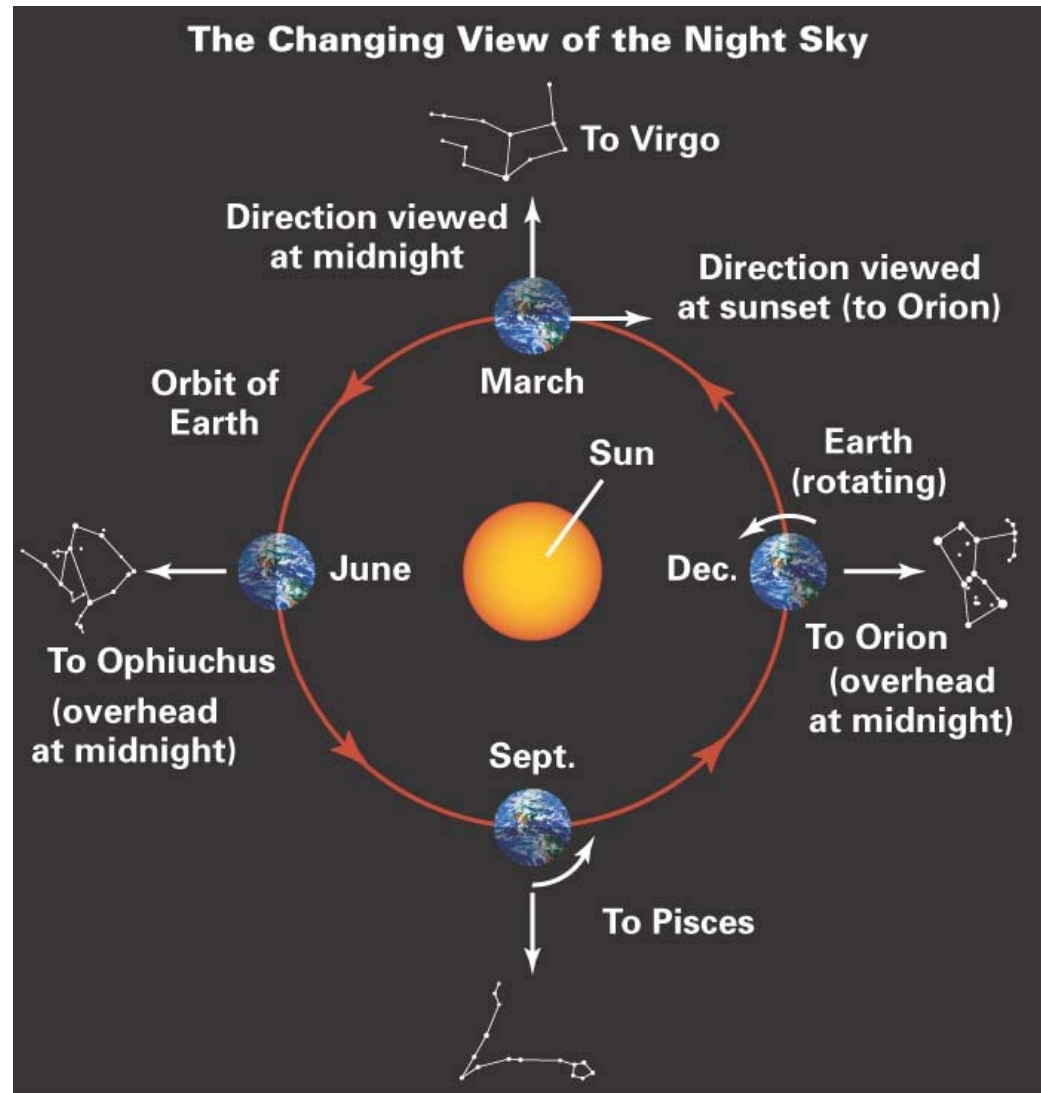
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# The Earth Revolves Around the Sun

- The Earth revolves around the sun once per year
- The distance from the Sun to the Earth changes by about 3% over a year





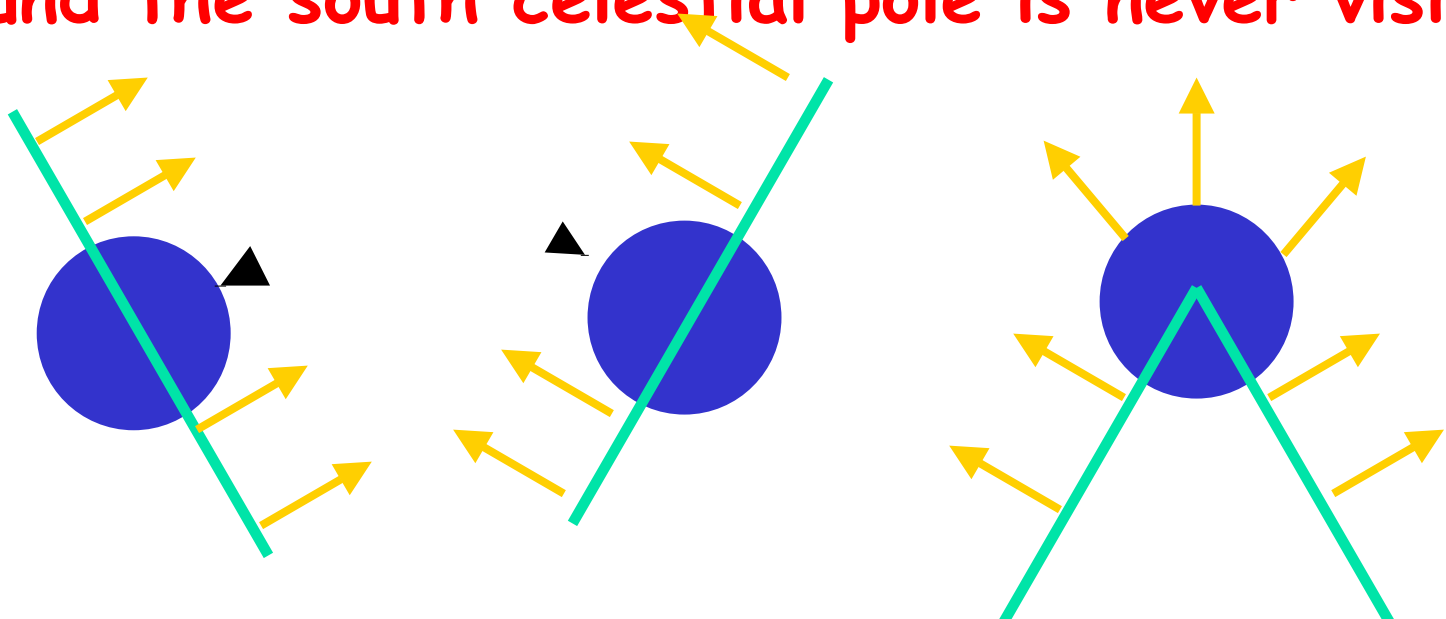
# What Can I See?

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- Earth's rotation and motion around the Sun determine what we can see in the sky
- Rotation
  - During the day: the Sun
  - At night: stars "rising" in the east and "setting" in the west
- Motion around the Sun
  - Six months from now the current sky will be hidden by the Sun and we will see part which is now behind the Sun

# What Can I See? (cont.)

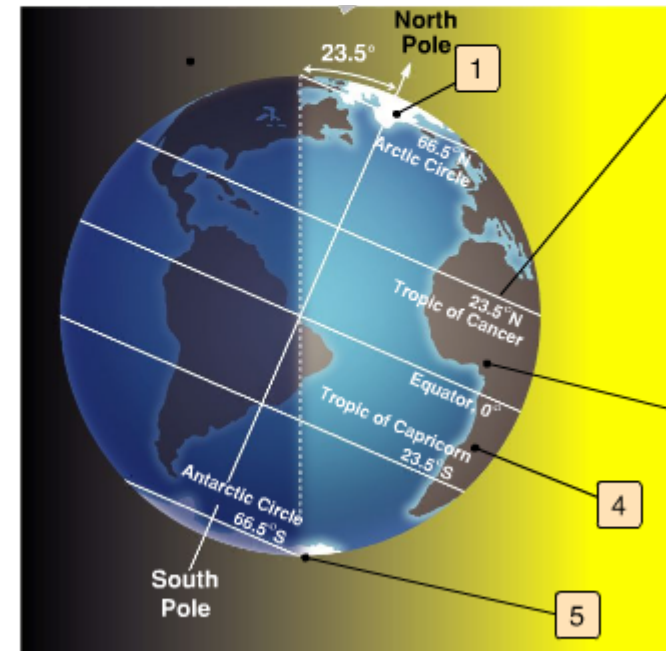
- In the northern hemisphere, the North Star is above the horizon all day long
  - The angle of the North Star above the horizon equals your latitude
- In the northern hemisphere, part of the sky around the south celestial pole is never visible





# The Earth is Tilted

- The Earth's axis is tilted  $23.5^\circ$  with respect to its orbit around the Sun
  - Axis always points in the same direction, toward the north celestial pole
    - It actually moves very slowly over time, precessing like a top



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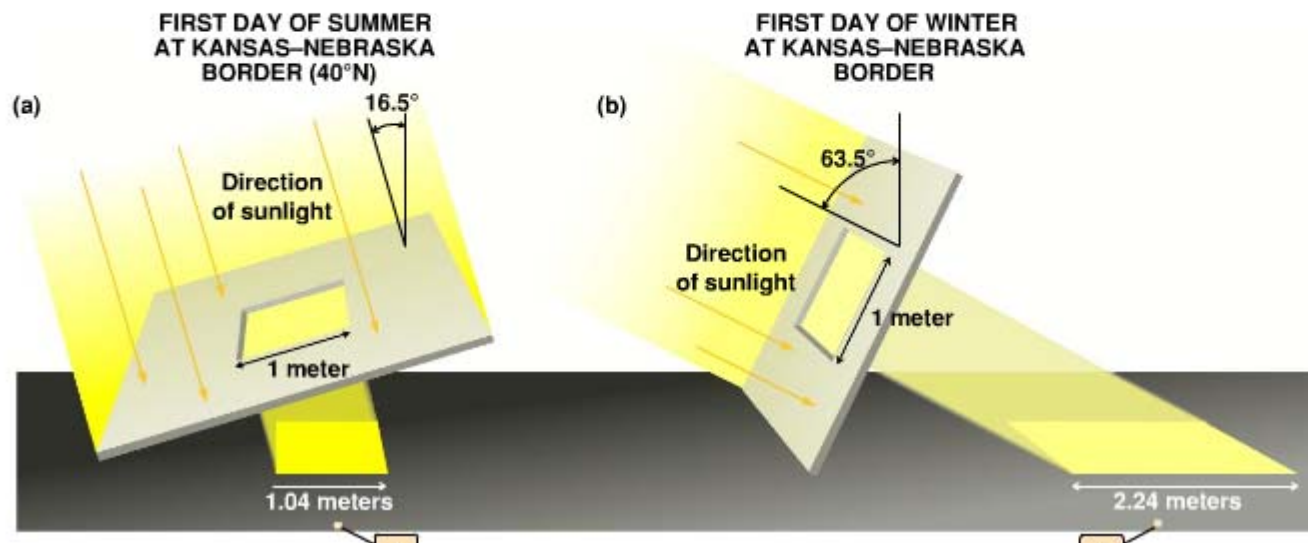
# Seasons

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- Seasons are caused by the tilt of the Earth combined with motion around the Sun
- During our summer, the north celestial pole is pointed towards the Sun
  - The Sun is above the horizon longer
  - We receive more intense light
- During our winter, the south celestial pole is pointed towards the Sun

# Intensity of Sunlight

- The amount of light per square meter depends on the angle at which the light hits the surface
- The amount of light determines the "heating" of the Earth
- In the summer, the light is more direct





# Summer Solstice

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- **First day of summer, about June 22**
  - Sun appears to be  $23^\circ$  north of the equator
    - Passes through the zenith of places that are  $23^\circ$  N latitude
    - $23^\circ$  N latitude is called the Tropic of Cancer
  - All regions within  $23^\circ$  of the North Pole see the sun for the full day
    - $90^\circ - 23^\circ = 67^\circ$  N latitude is called the Arctic Circle
  - All regions within  $23^\circ$  of the South Pole see no sunlight for the full day
    - $67^\circ$  S latitude is called the Antarctic Circle
    - It stays dark at the North and South Poles for 6 months each year



# Winter Solstice and Equinoxes

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- **First day of winter, about December 22**
  - Everything is reversed
  - Sun passes through the zenith along the Tropic of Capricorn at noon ( $23^{\circ}$  S latitude)
- **Equinoxes**
  - Twice a year, the sun passes through the zenith along the equator at noon ( $0^{\circ}$  latitude)
  - Vernal Equinox, around March 21
  - Autumnal Equinox, around September 21
  - 12 hours of light and 12 hours of darkness everywhere



# Length of the Year

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- It takes the Earth 365.242199 days to go from one vernal equinox to the next
  - NOT an integer number
  - But the extra is close to  $\frac{1}{4} = 0.25$
- So every 4 years (leap year) we add an extra day to the calendar (Feb. 29)
  - But this is too much (we've added 0.25!)
- So every 100 years (on the century) we don't add the extra day (no leap year)
- But this isn't right either, so every 4<sup>th</sup> 100 years, we do include the leap year
- This is why 2000 was a leap year



# Summary

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- The Universe has lots of motion
- The spinning of the Earth causes the rising and setting of the Sun and stars
- The revolution of the Earth around the Sun determines the year
- The tilt of the Earth determines the seasons
- The spinning, revolution and tilt determine the part of the sky which is visible
- You want/need to understand these motions
- Next time, we will look at how the Moon behaves



# Real World

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- Earth's atmosphere fuzzes the edges (figuratively and literally)
- Atmosphere bends light coming from the Sun, allowing us to "see over the horizon" about  $18^\circ$ 
  - Sun appears to rise earlier and set later
  - It's light out (twilight) in the morning when the Sun is  $18^\circ$  below the horizon and stays light in the evening until the Sun is  $18^\circ$  below the horizon
  - Effect is most noticeable at the poles - complete darkness for only 3 months (rather than 6 months)
    - Last week the scientific station in Antarctica starting receiving flights after the winter break
- June 22 is the longest day, but not the hottest, why?